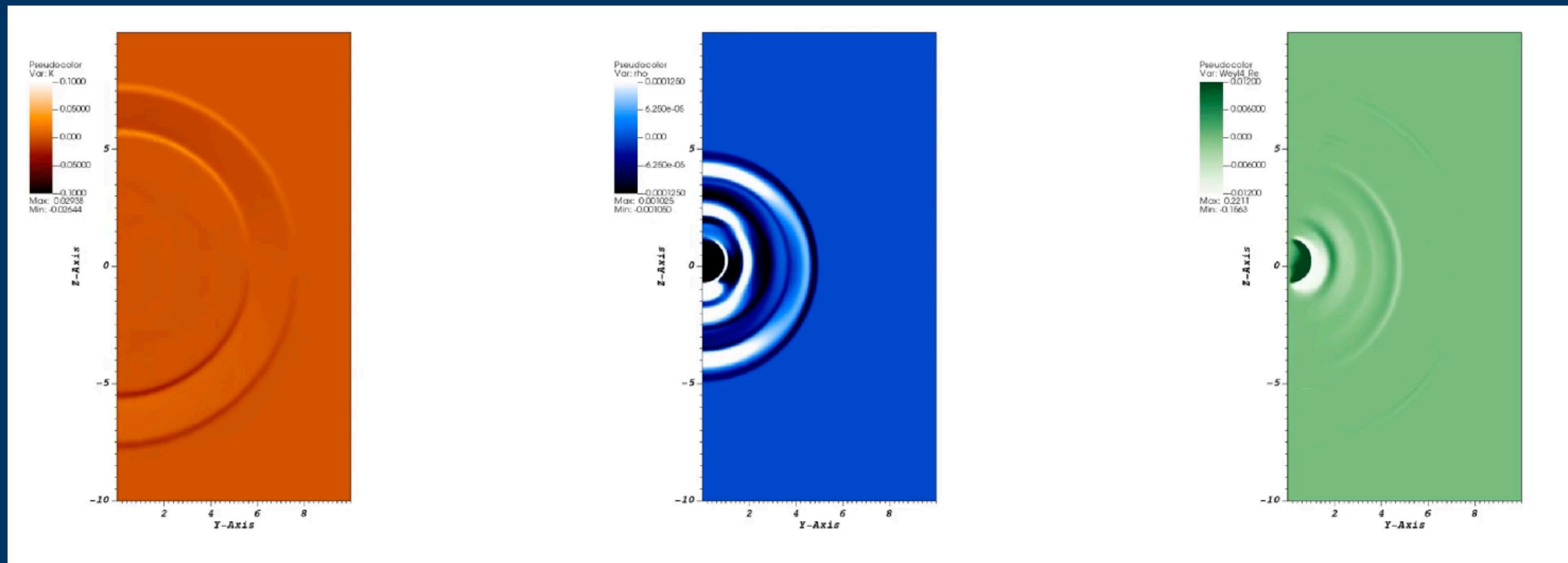


What no one has seen before: gravitational waveforms from warp drive collapse

Katy Clough

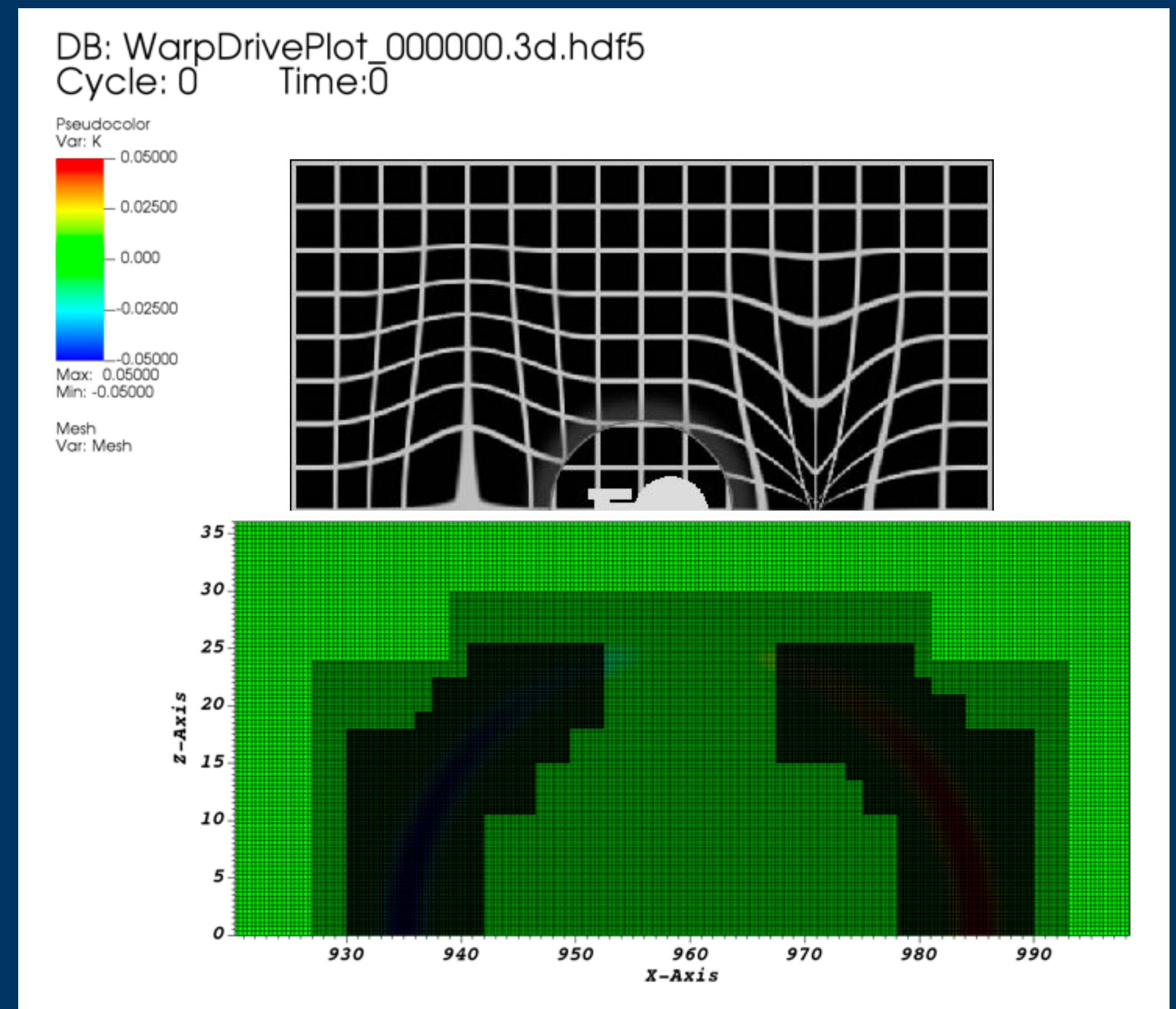
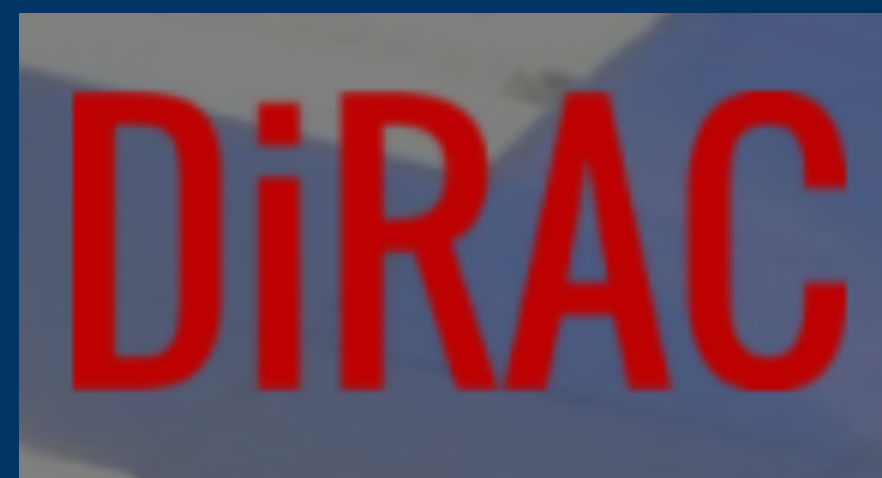


Last time on ETK, the next generation...

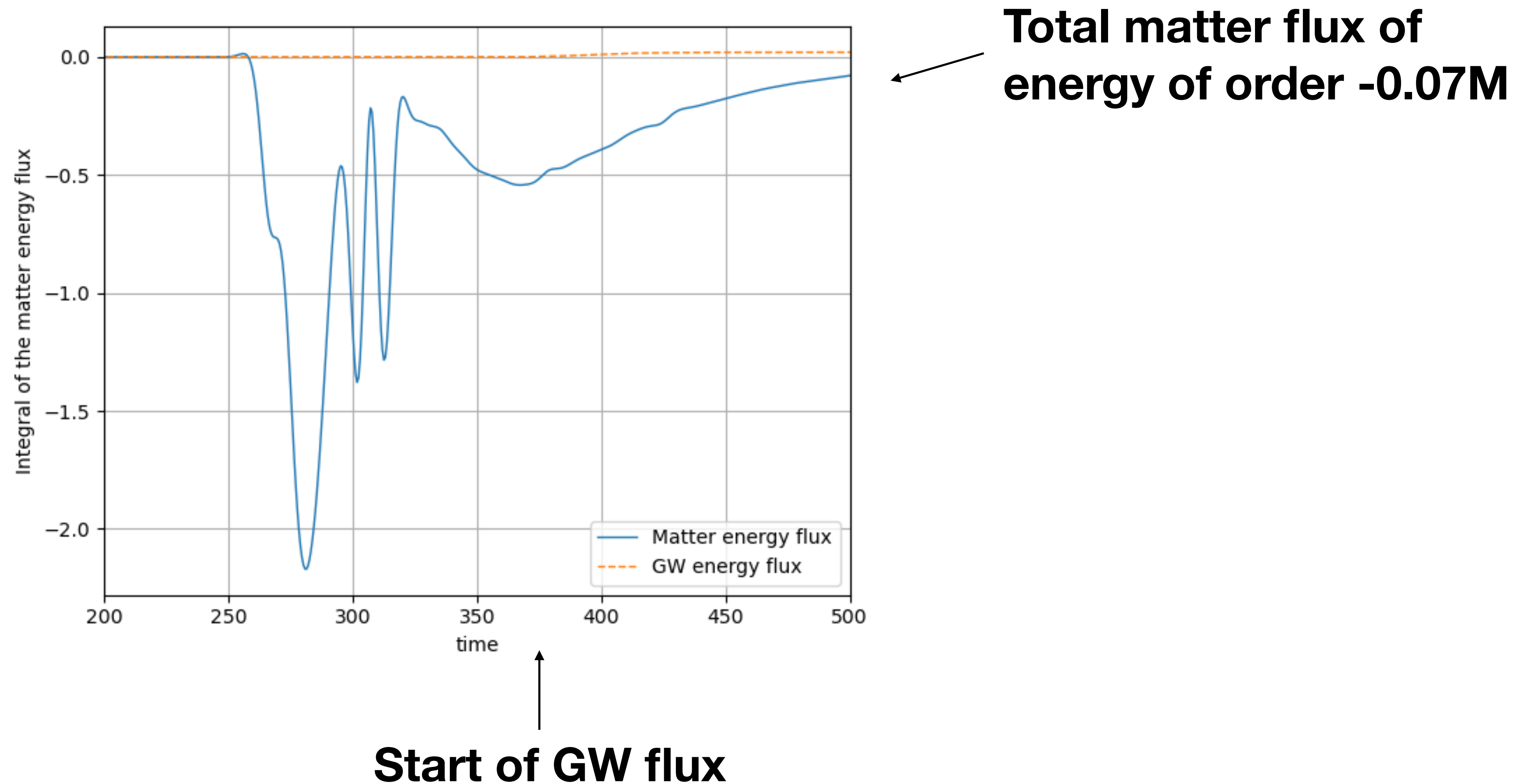


Adaptive mesh refinement (and warp drives)

Katy Clough



Some rather dodgy looking plots followed...



Conclusion:

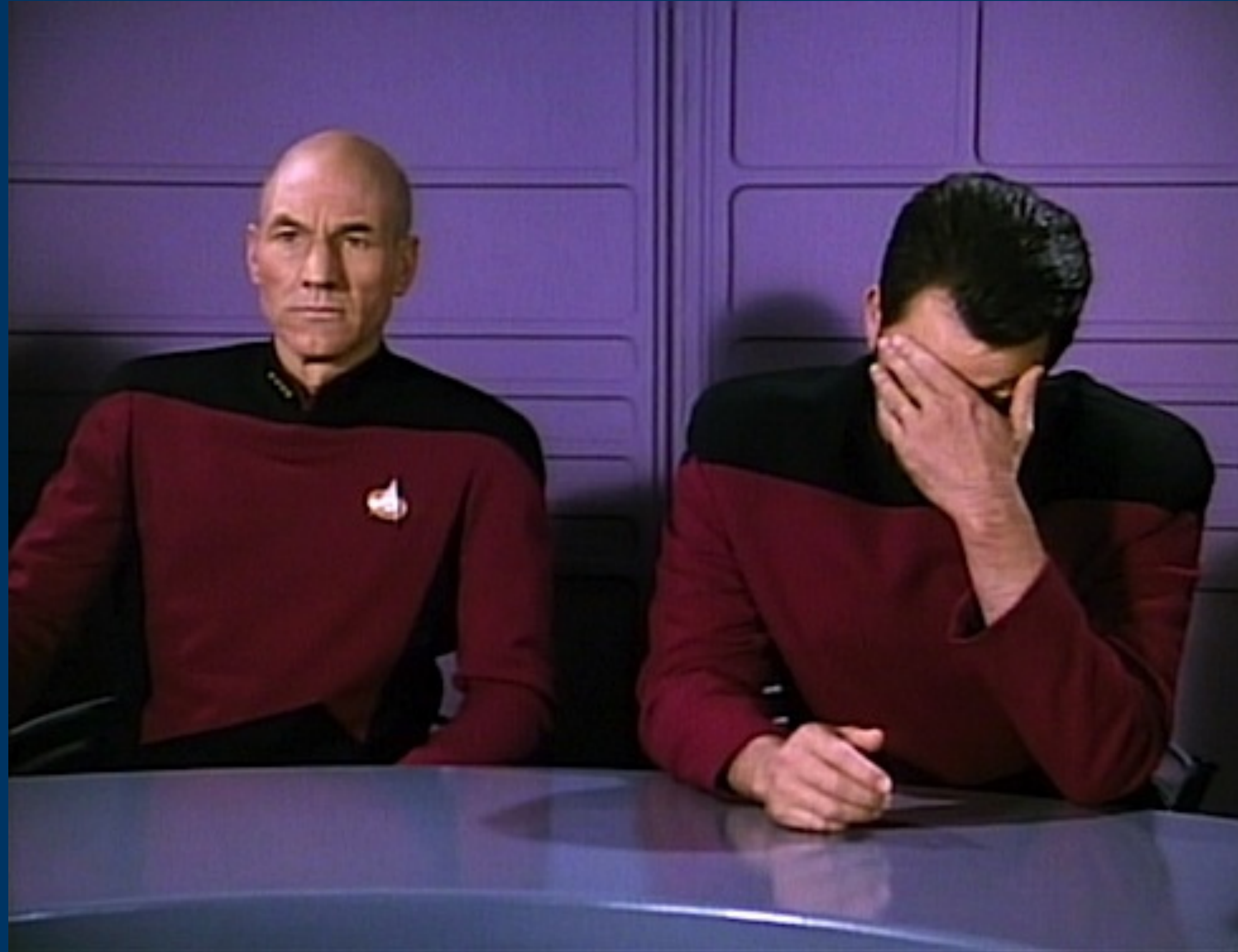
Still many questions...

Not sure it is actually working...

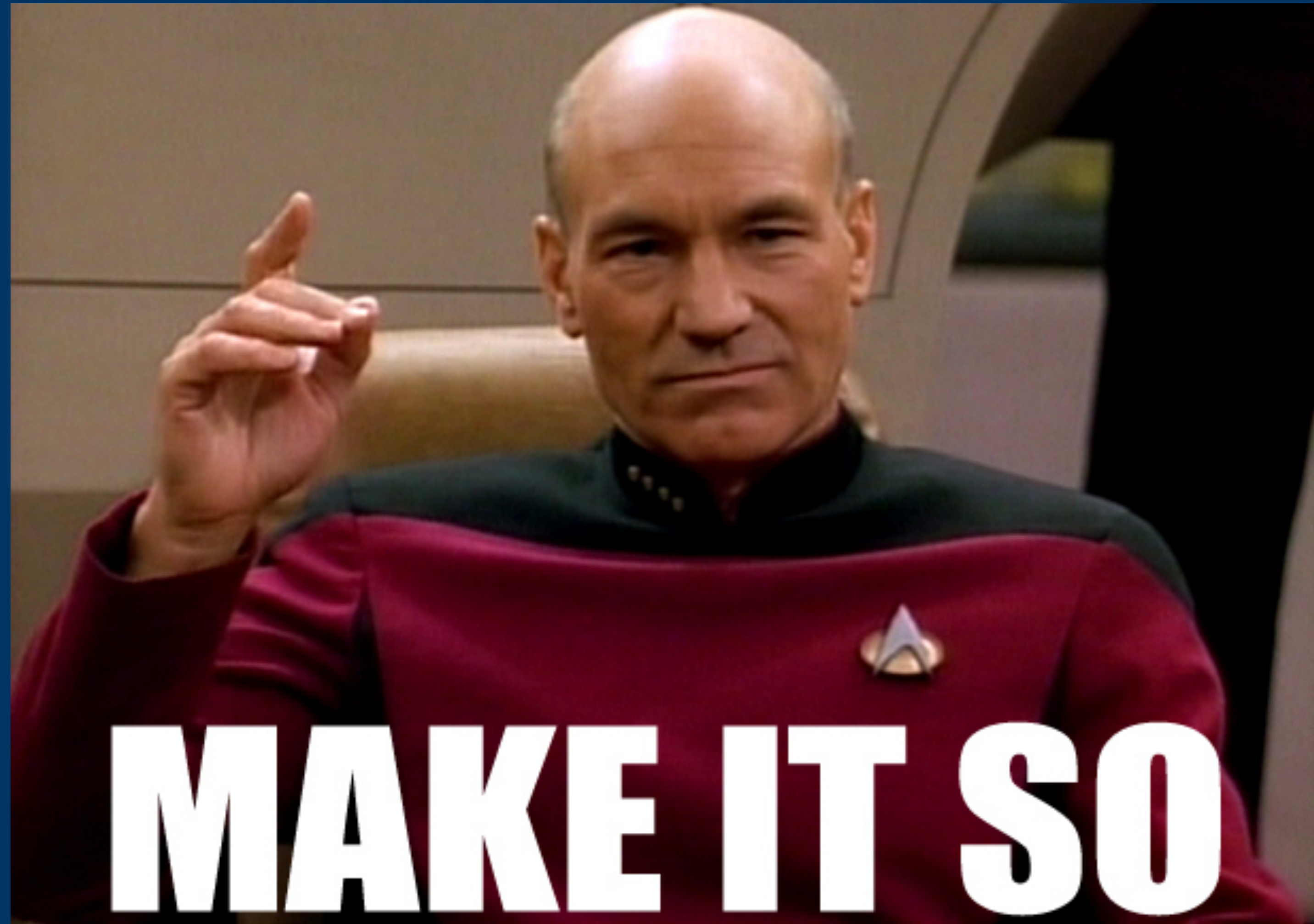
Does this even make sense?



Other conferences were like...



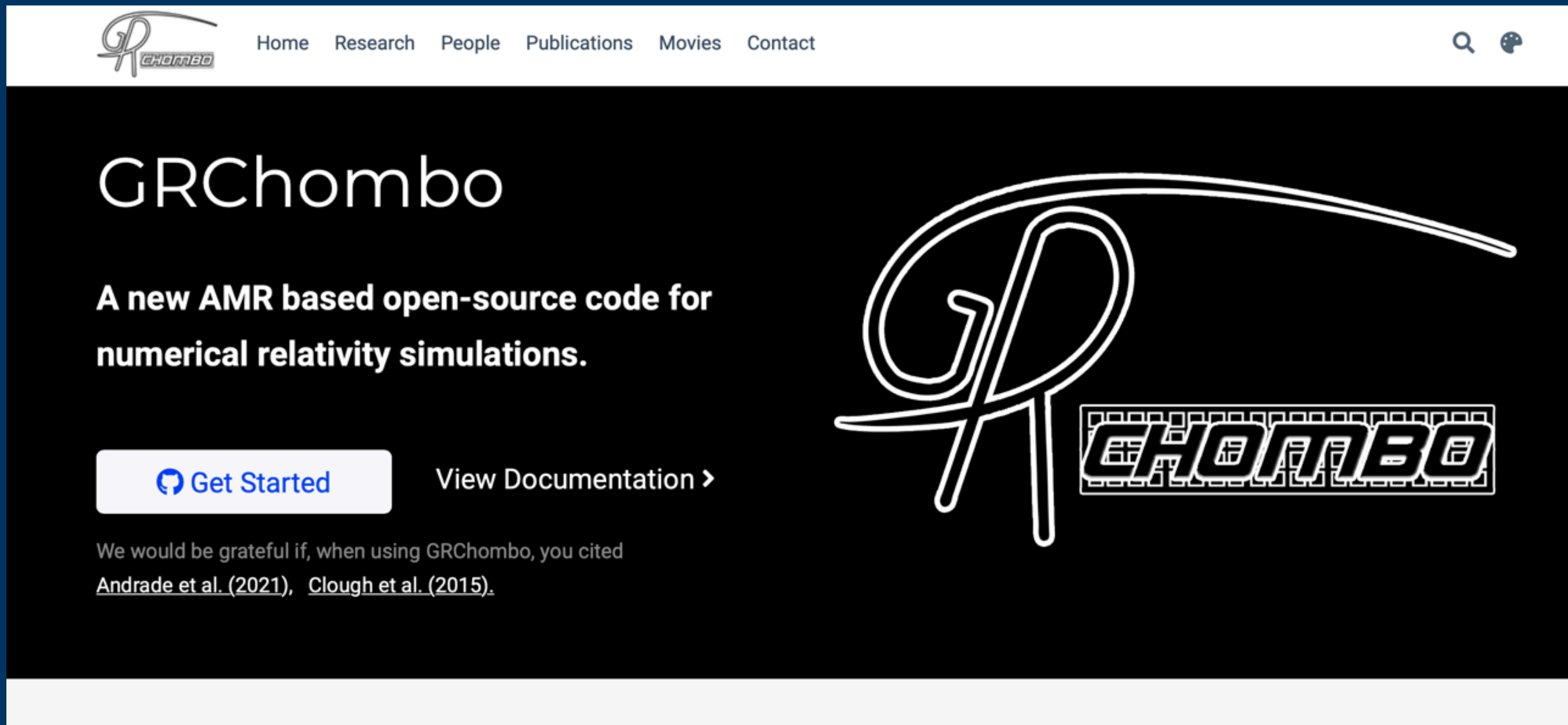
ETK were like...



And now the conclusion...



About me: a friend of ETK



The screenshot shows the GRChombo website with a white header containing navigation links and a search icon. The main content area has a black background with the project name, a description, and call-to-action buttons. A large logo is on the right, and a citation notice is at the bottom.

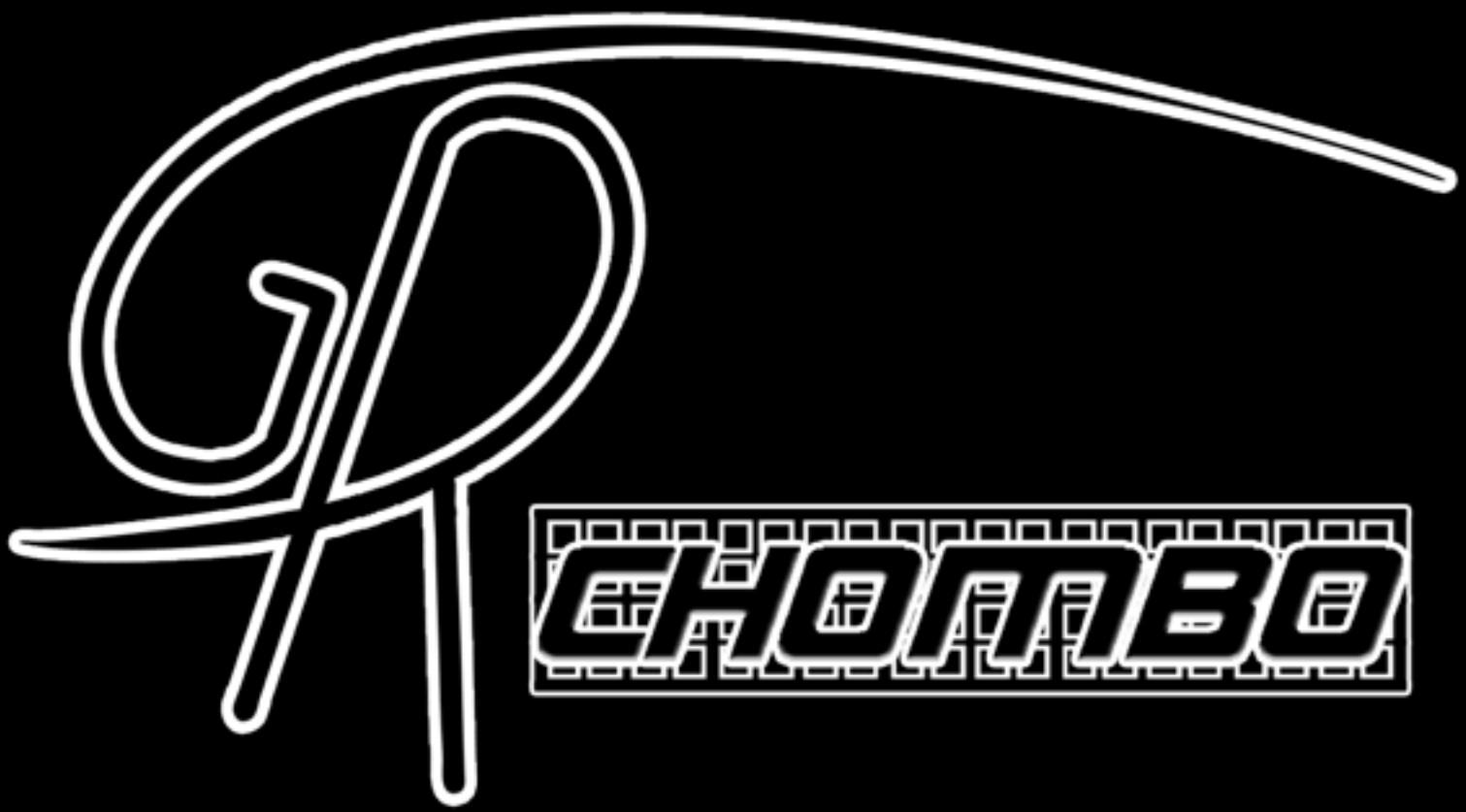
GRChombo

Home Research People Publications Movies Contact

A new AMR based open-source code for numerical relativity simulations.

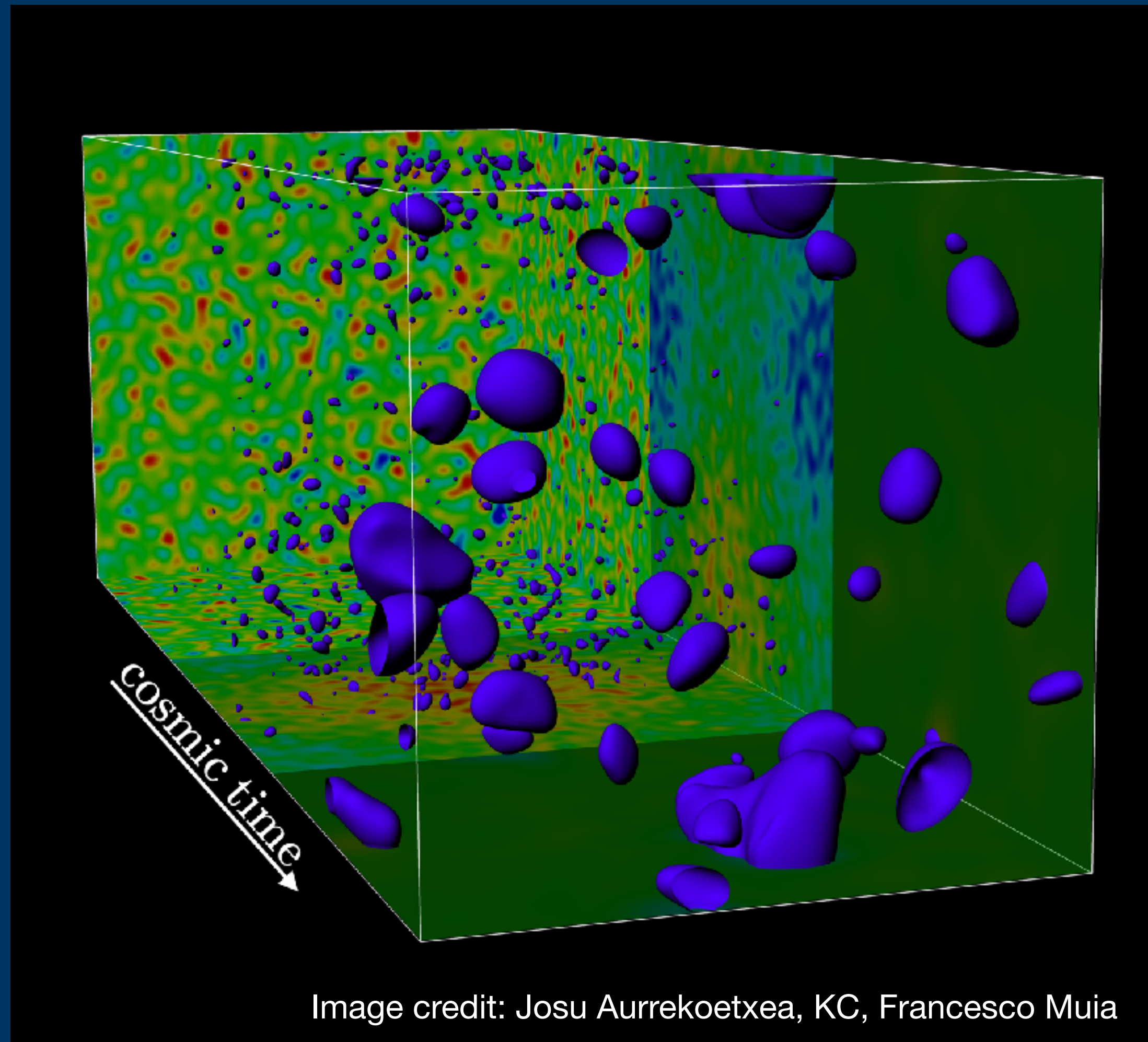
[Get Started](#) [View Documentation >](#)

We would be grateful if, when using GRChombo, you cited [Andrade et al. \(2021\)](#), [Clough et al. \(2015\)](#).



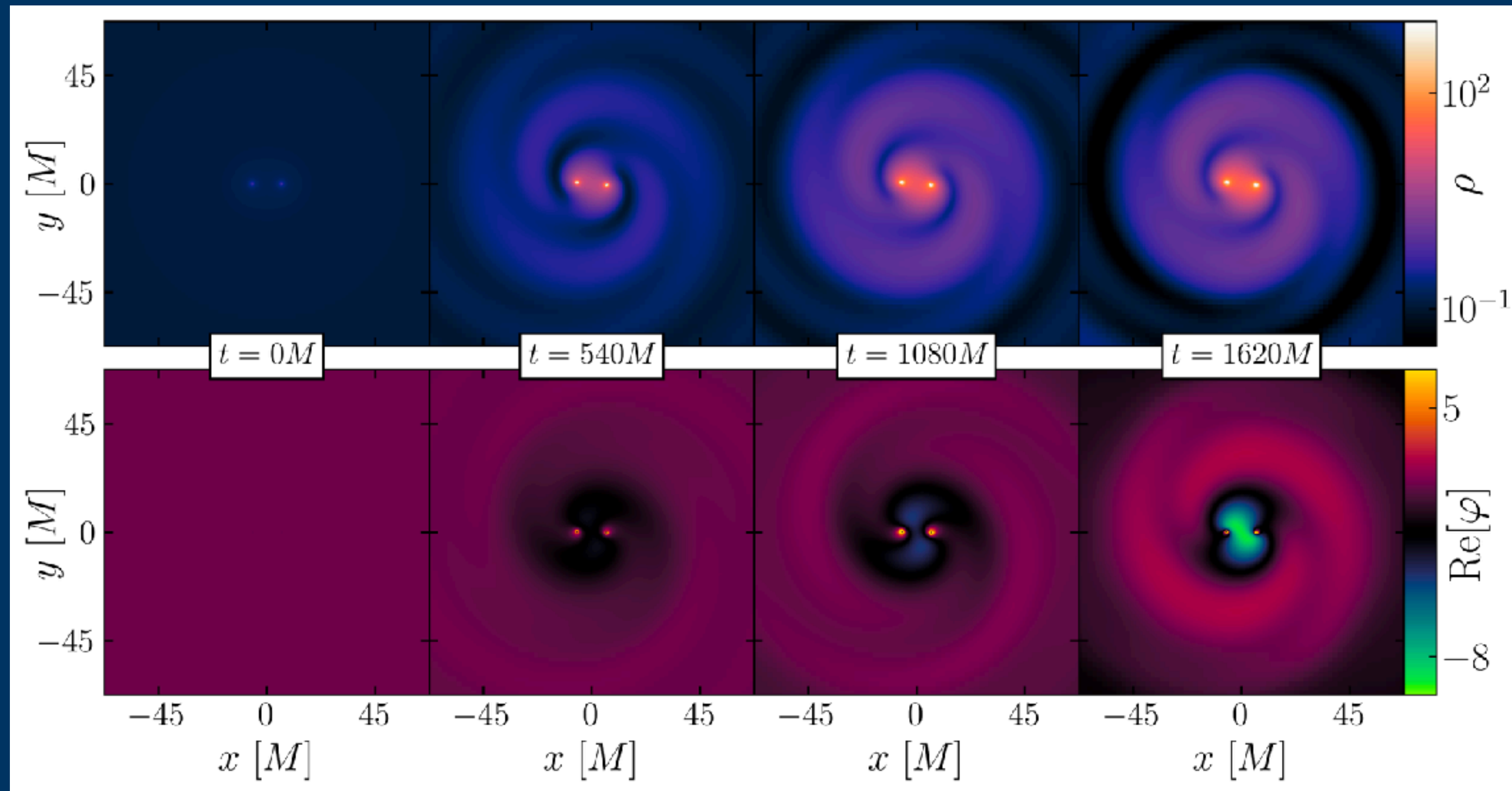
The logo features a large, stylized 'GR' in a white outline font, with the word 'CHOMBO' in a white, blocky, outlined font to its right.

I do have some more serious projects:



Inhomogeneous inflation

I do have some more serious projects:



Fundamental
fields and BHs

But today: A fun(damental) example

Preview Abstract

What no one has seen before: gravitational waveforms from warp drive collapse

Katy Clough, Tim Dietrich and Sebastian Khan

Despite originating in science fiction, warp drives have a concrete description in general relativity, with Alcubierre first proposing a spacetime metric that supported faster-than-light travel. Whilst there are numerous practical barriers to their implementation in real life, including a requirement for negative energy, computationally, one can simulate their evolution in time given an equation of state describing the matter. In this work, we study the signatures arising from a warp drive "containment failure", assuming a stiff equation of state for the fluid. We compute the emitted gravitational-wave signal and track the energy fluxes of the fluid. Apart from its rather speculative application to the search for extraterrestrial life in gravitational-wave detector data, this work is interesting as a study of the dynamical evolution and stability of spacetimes that violate the null energy condition. Our work highlights the importance of exploring strange new spacetimes, to (boldly) simulate what no one has seen before.

Comments: 12 pages, 6 figures, plus appendix. Comments welcome!

License: <http://arxiv.org/licenses/nonexclusive-distrib/1.0/>

Categories

The crew

Katy "I cannae do it Captain" Clough



Captain Tim "Jim T." Dietrich



Sebastian "Wrath of" Khan

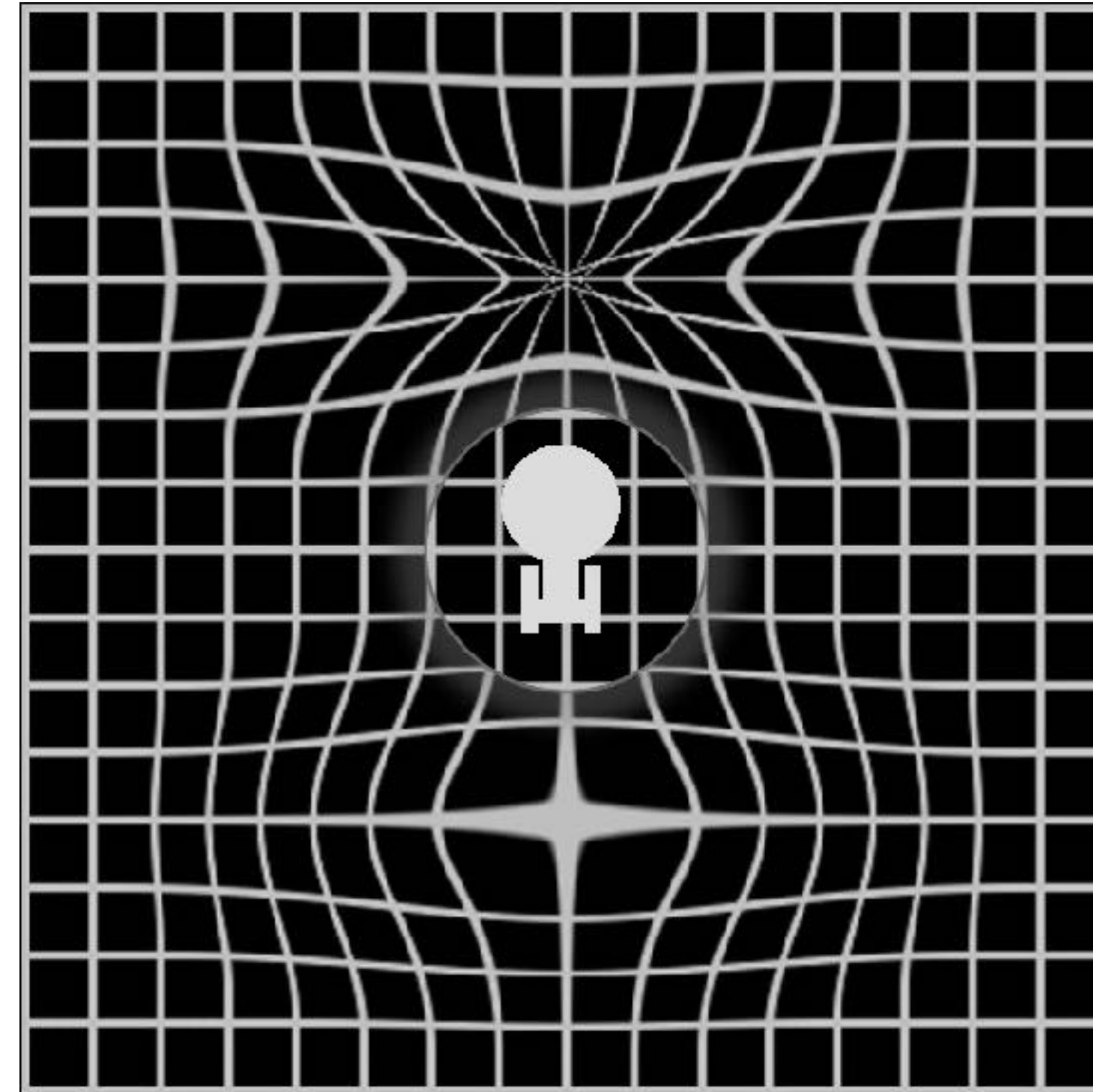
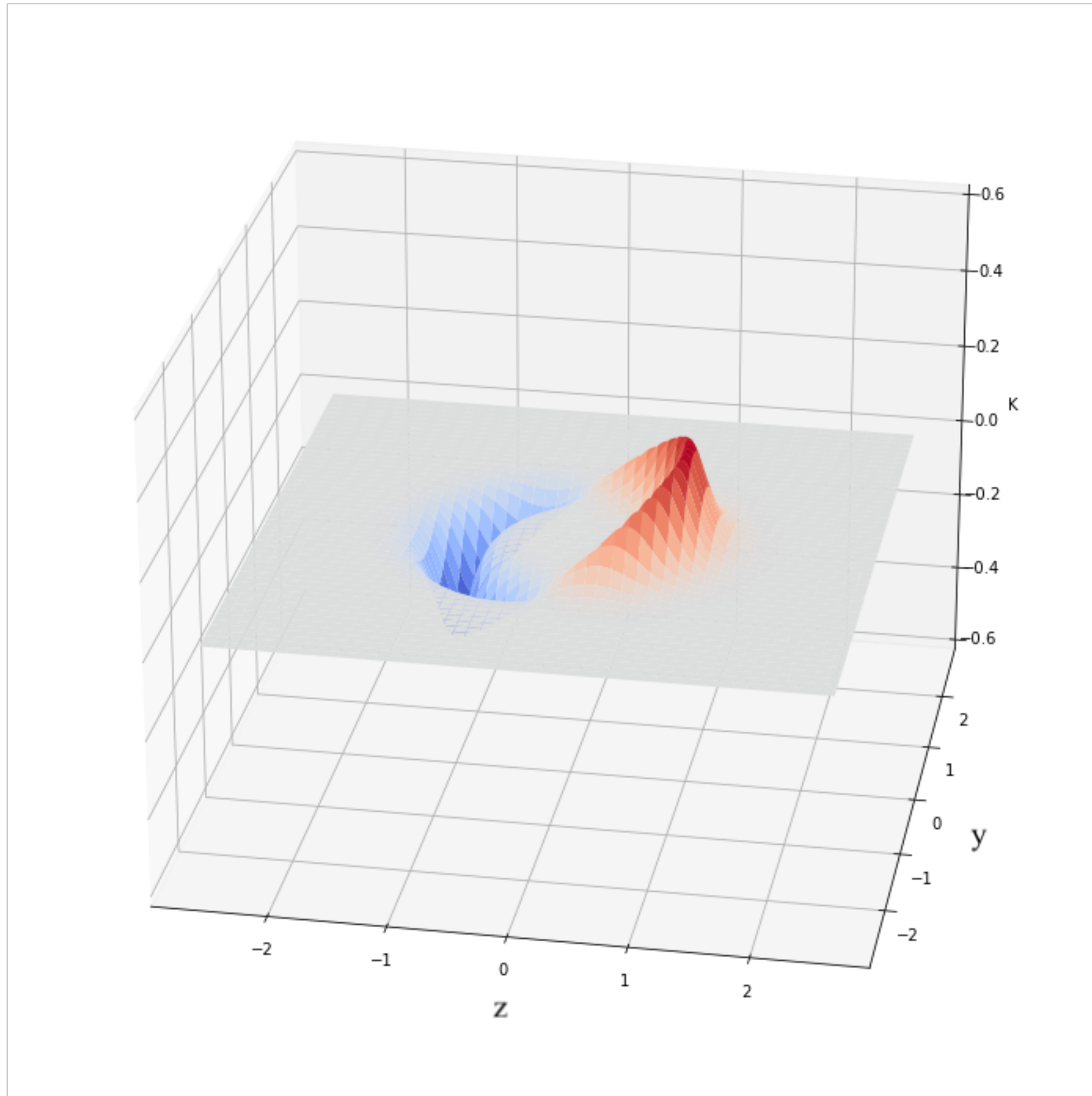


**Question 1: Could we
see gravitational
waves from alien warp
drive ships?**



**Question 2: Can
asymptotically zero
ADM mass spacetimes
radiate?**

A warp drive metric exists



Alcubierre, The warp drive: hyper-fast travel within general relativity, CQG 1994

Tasks for the chief engineer:

- Initial conditions - for metric and fluid
- Evolution equations - for metric and fluid
- Specify coordinates / gauge
- Diagnostics - what to measure?

Initial conditions for the spacetime metric

$$\gamma_{ij} = \delta_{ij}$$

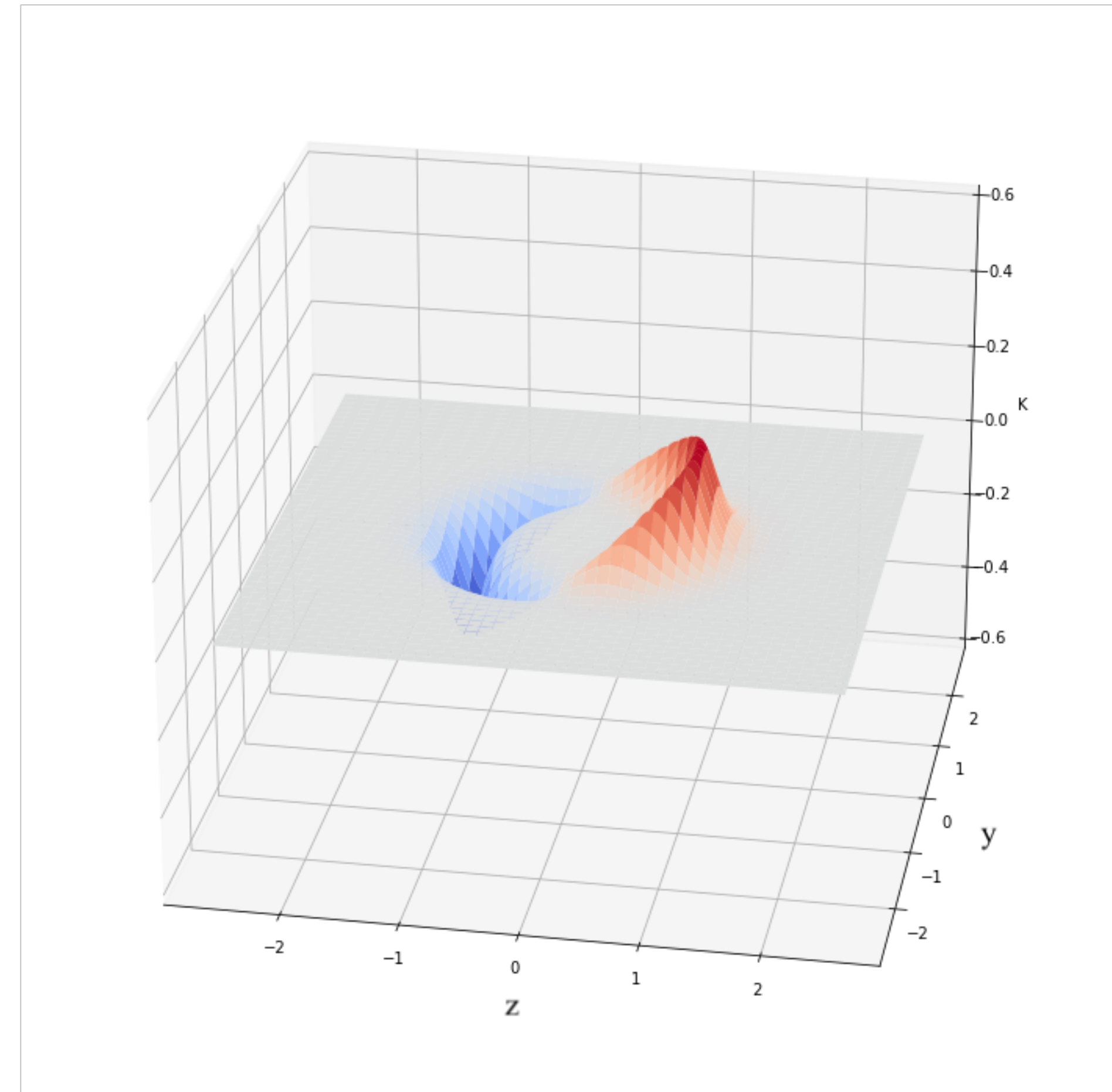
$$\alpha = 1$$

$$\beta^z = f(r - R)$$

$$K = \partial_z \beta^z$$

This requires matter that violates the
Null Energy Condition,

(ie negative energy density)



Initial configuration for the matter

$$T_{\mu\nu}(t = 0) = G_{\mu\nu}^{alcubierre}(t = 0)/(8\pi)$$

This requires matter that violates the
Null Energy Condition,
(ie negative energy density)



Metric evolution - just GR

The Einstein equation

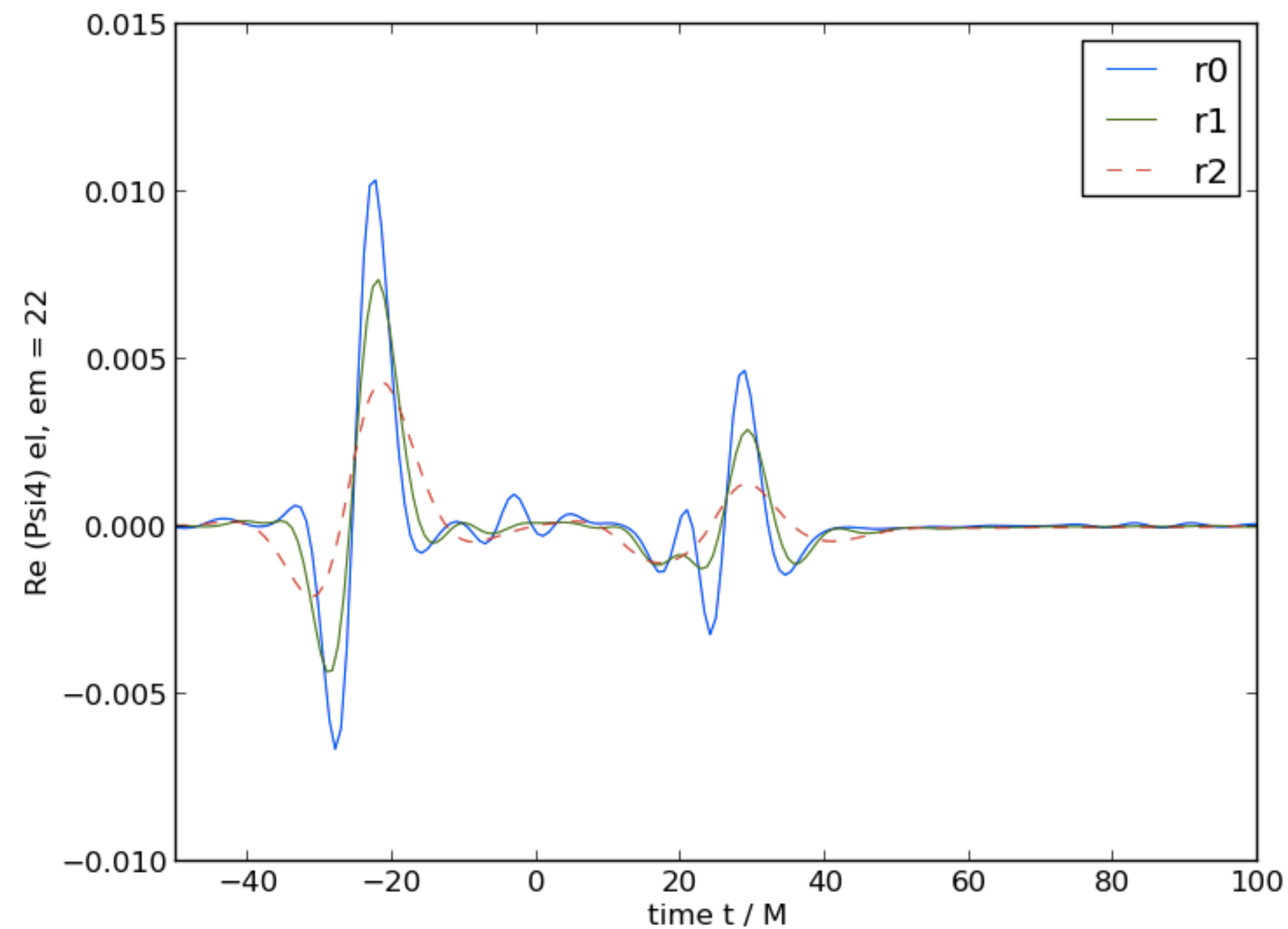
$$R_{ab} - R/2 g_{ab} = 8\pi T_{ab}$$

t

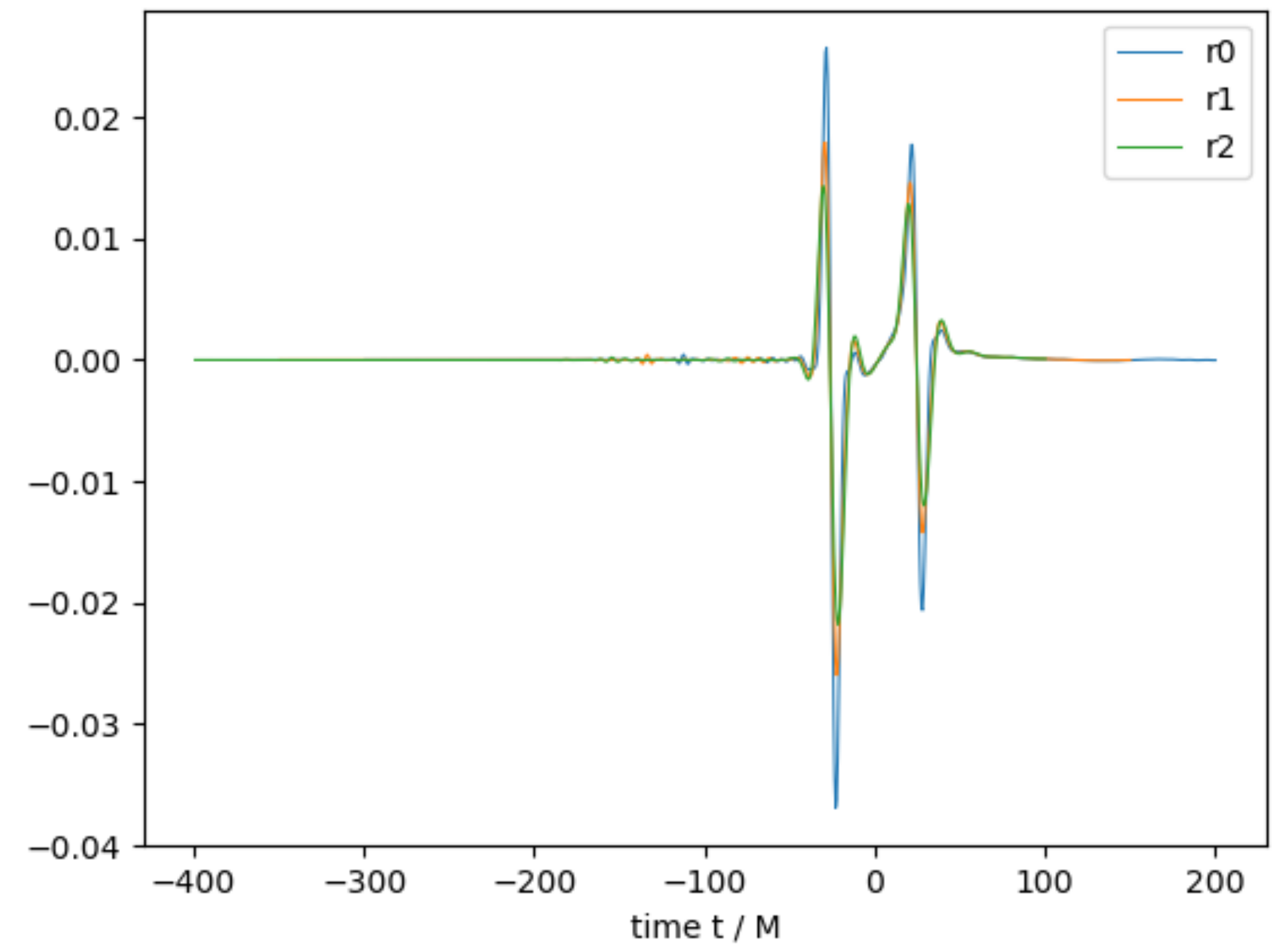
$f(\partial^2 g_{ab}, \partial g_{ab}, g_{ab})$
“Curvature”

“Energy-Momentum”

But finding good coordinates is a challenge



Bad gauge choice



Better gauge choice

But finding good coordinates is a challenge

Shock-avoiding slicing conditions: Tests and calibrations

Thomas W. Baumgarte and David Hilditch
Phys. Rev. D **106**, 044014 – Published 4 August 2022

[Article](#)[References](#)[Citing Articles \(4\)](#)[PDF](#)[HTML](#)[Export Citation](#)

ABSTRACT

While the $1 + \log$ slicing condition has been extremely successful in numerous numerical relativity simulations, it is also known to develop “gauge shocks” in some examples. Alternative “shock-avoiding” slicing conditions suggested by Alcubierre prevent these pathologies in those examples, but have not yet been explored and tested very broadly. In this paper we compare the performance of shock-avoiding slicing conditions with those of $1 + \log$ slicing for a number of “text-book” problems, including black holes and relativistic stars. While, in some simulations, the shock-avoiding slicing conditions feature some unusual properties and lead to more “gauge dynamics” than the $1 + \log$ slicing condition, we find that they perform quite similarly in terms of stability and accuracy, and hence provide a very viable alternative to $1 + \log$ slicing.

Matter evolution

Local energy conservation tells us how 4 components evolve:

$$\nabla^\mu T_{\mu\nu} = 0$$

$$\implies \partial_t \rho = f(\rho, S_i, S_{ij}, g_{\mu\nu}) \quad \text{energy density}$$

$$\implies \partial_t S_i = f(\rho, S_i, S_{ij}, g_{\mu\nu}) \quad \text{momentum density}$$



Matter evolution

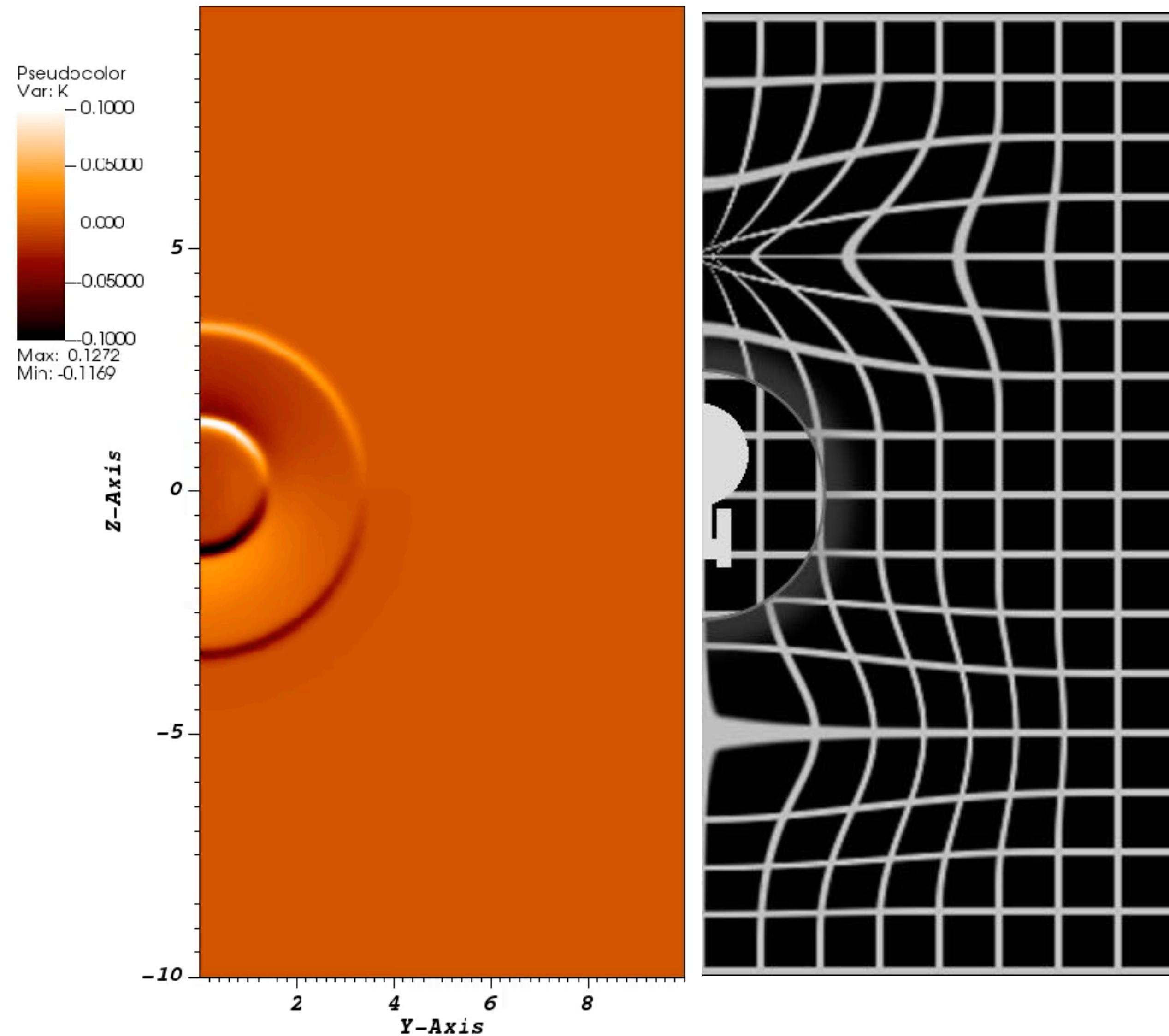
Still need eom for anisotropic stresses and pressures T_{ij} (effectively the fluid equation of state)

Bit of a hack:

Choose that anisotropic stresses decay in time, and a stiff equation of state $\bar{p} = \rho$.

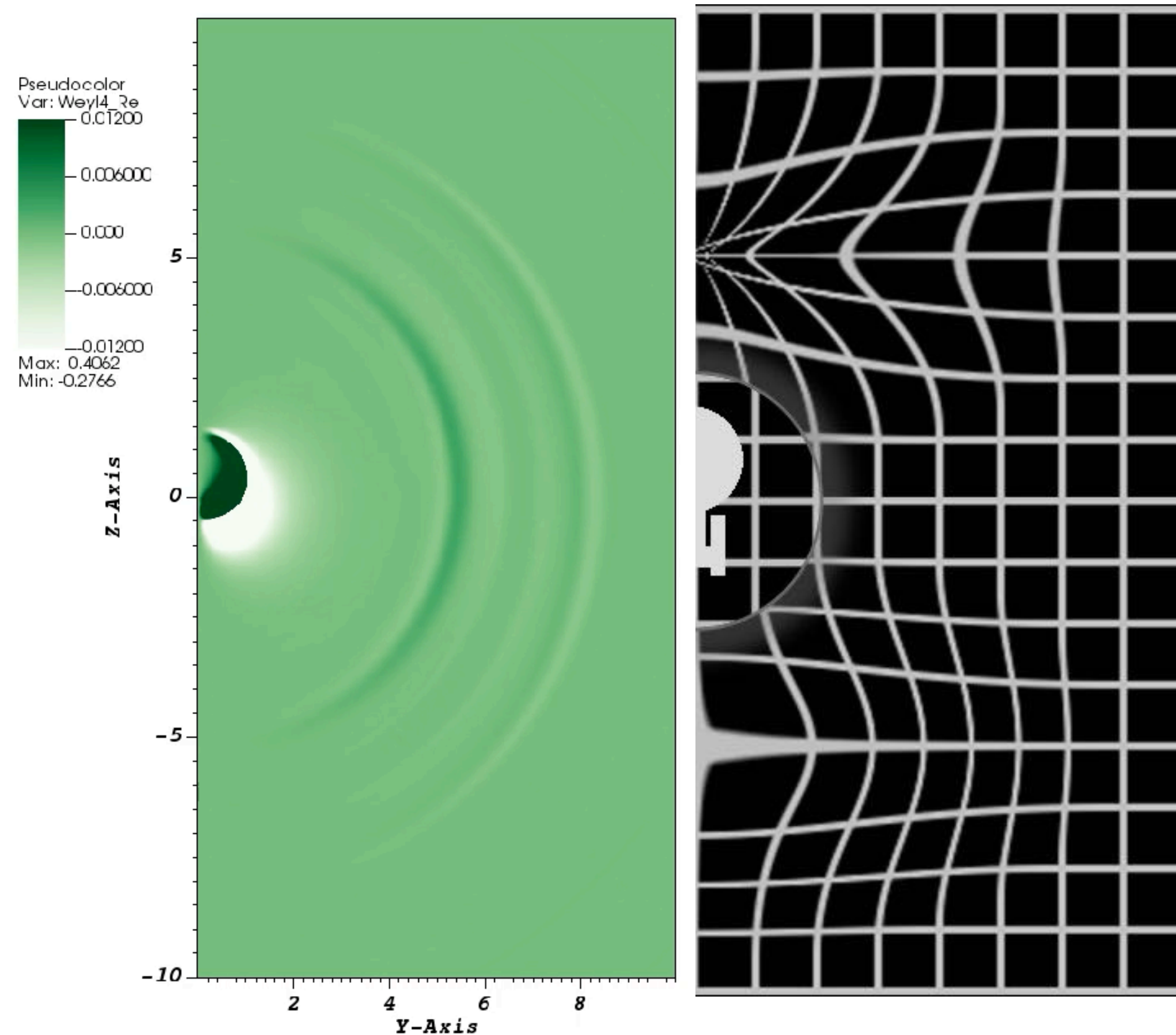


What no one has seen before...



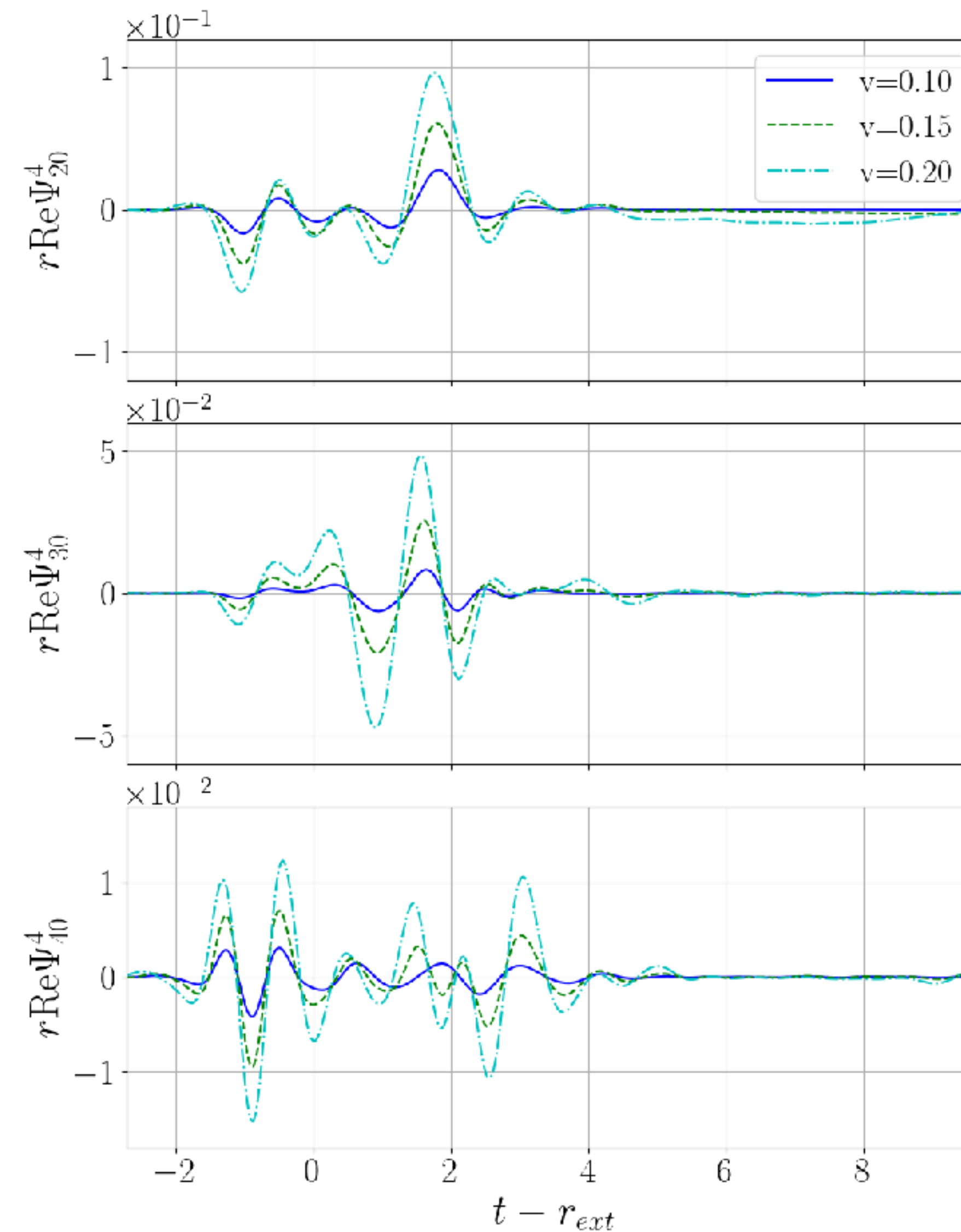
K
(Trace of extrinsic
curvature of spatial
slice)

One can extract the gravitational wave signals



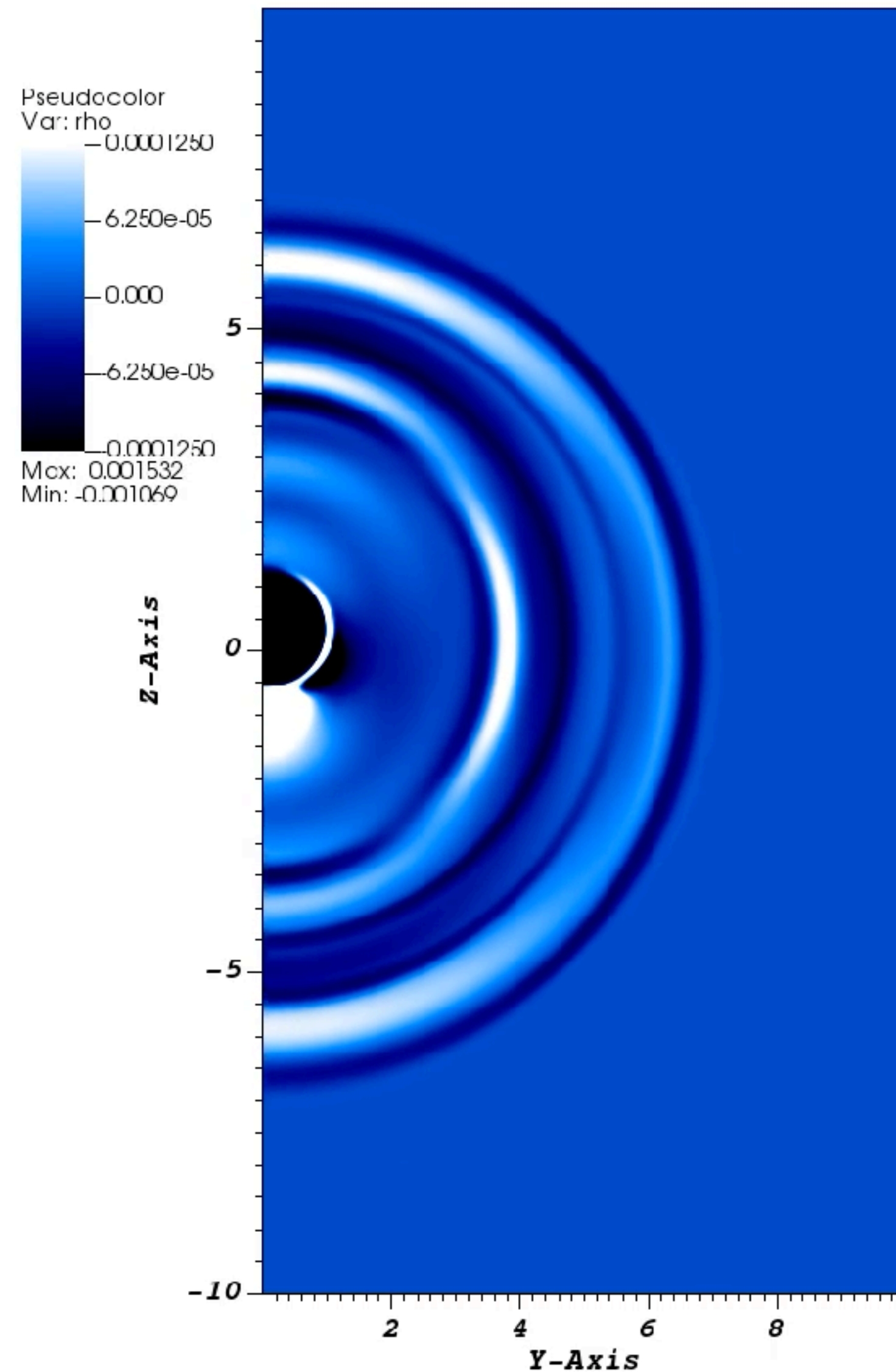
Weyl4 scalar

One can extract the gravitational wave signals



Weyl4 scalar

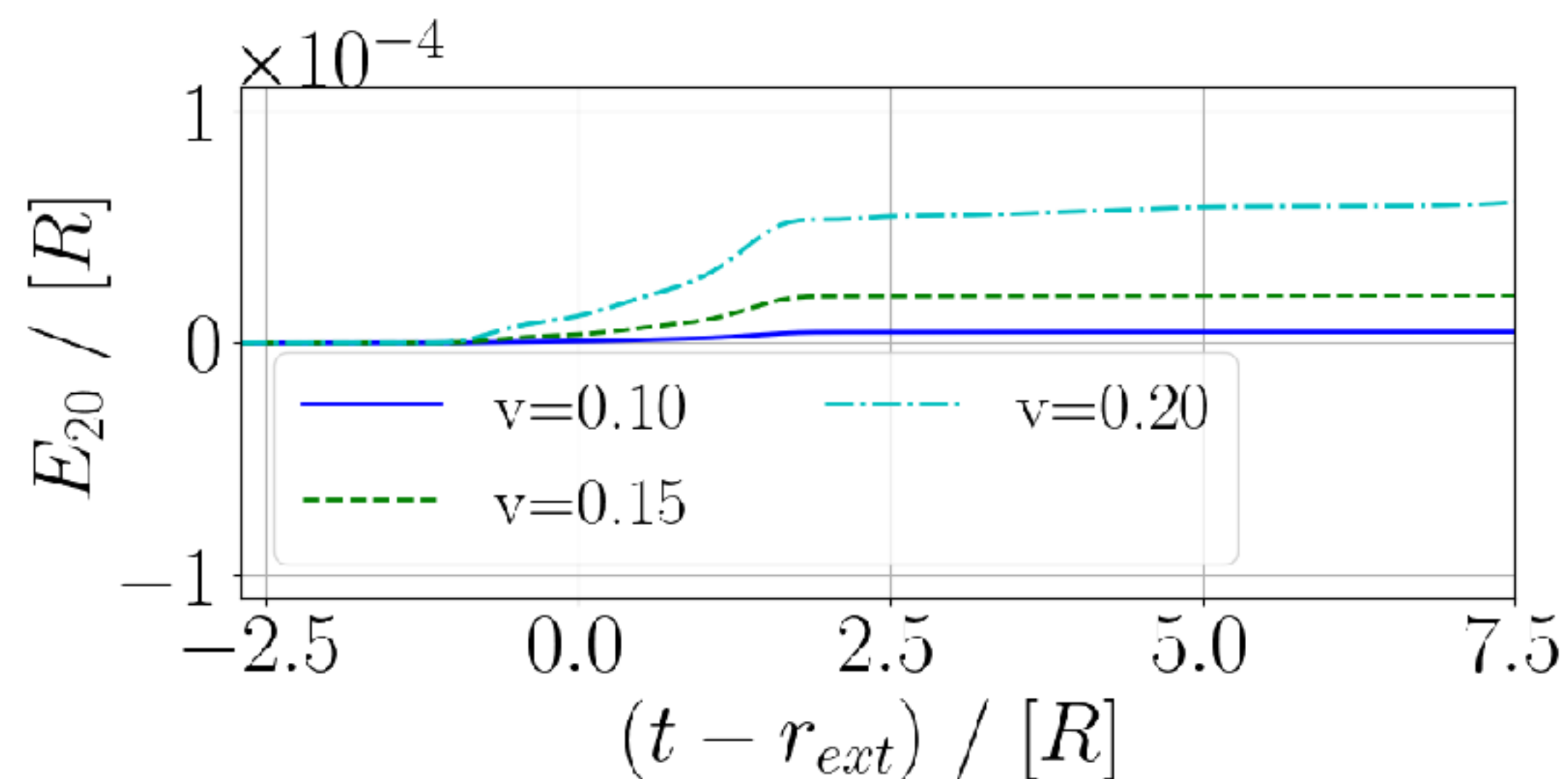
One can also study the fluxes of the matter content



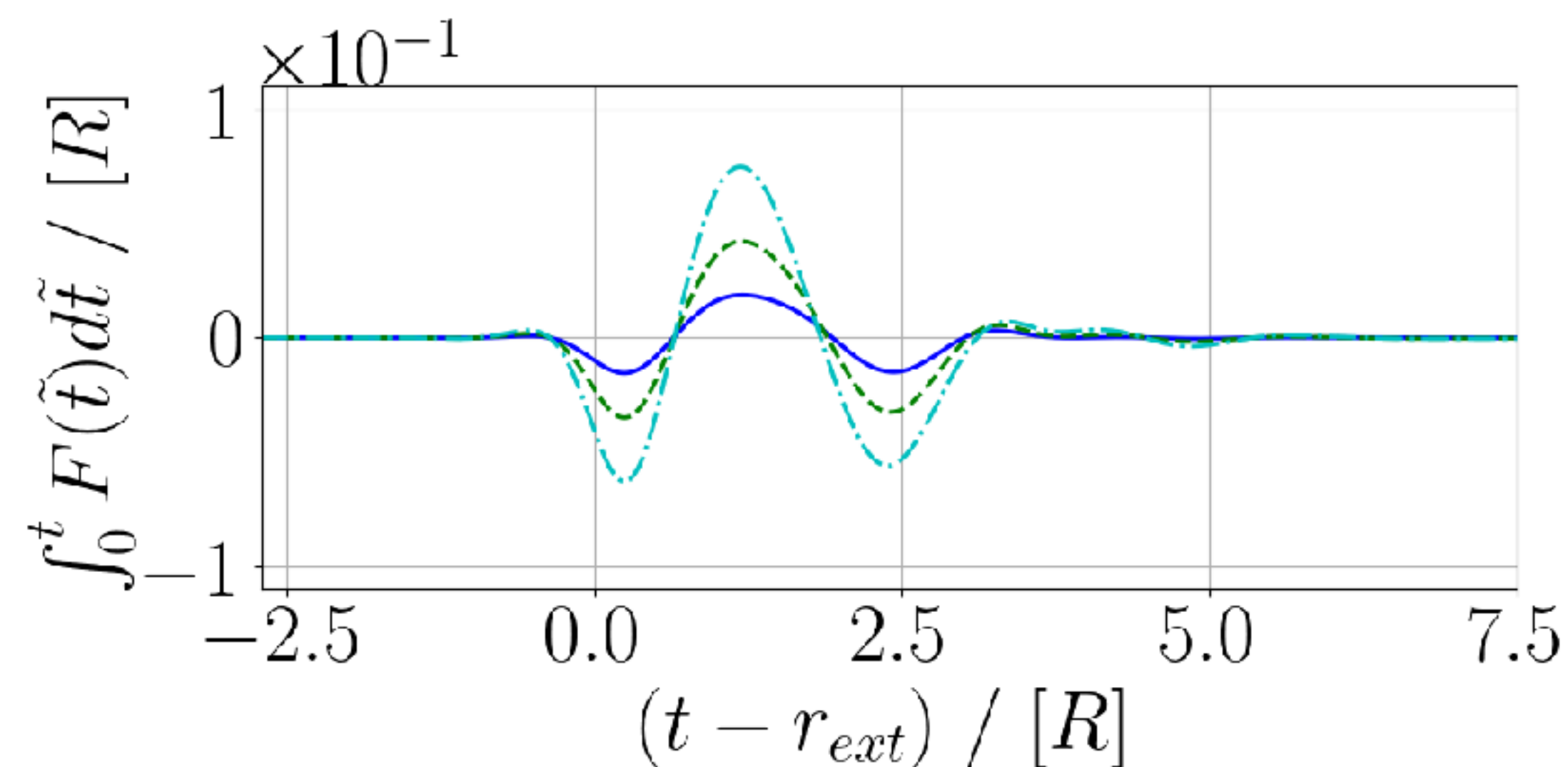
**Energy density
of warp fluid**

*See: Negative Mass
in General Relativity
(Bondi 1957)*

Matter energy radiation is both negative and positive

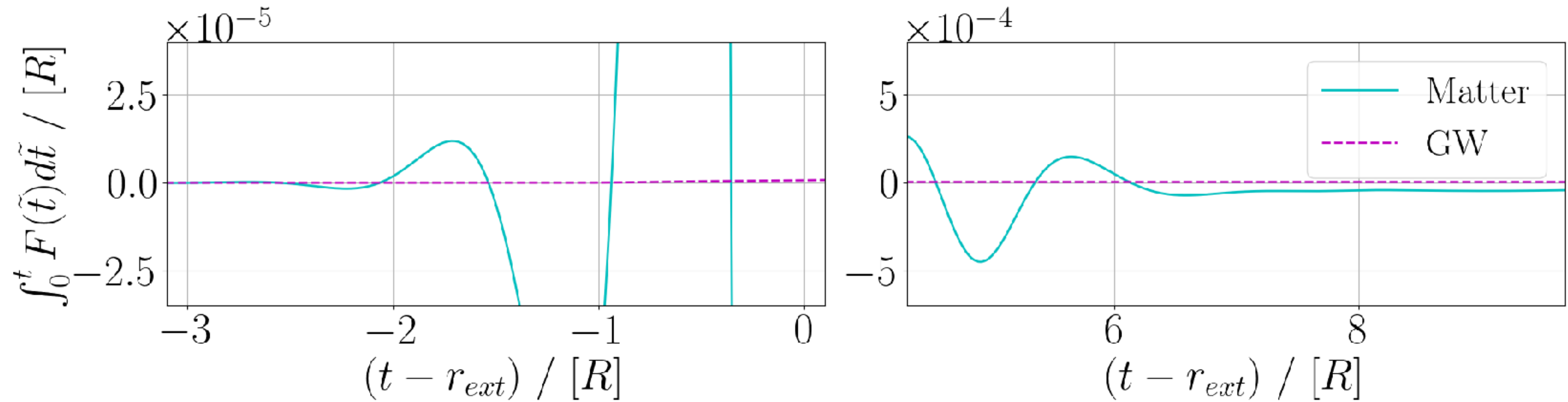


**Energy flux
of GWs**



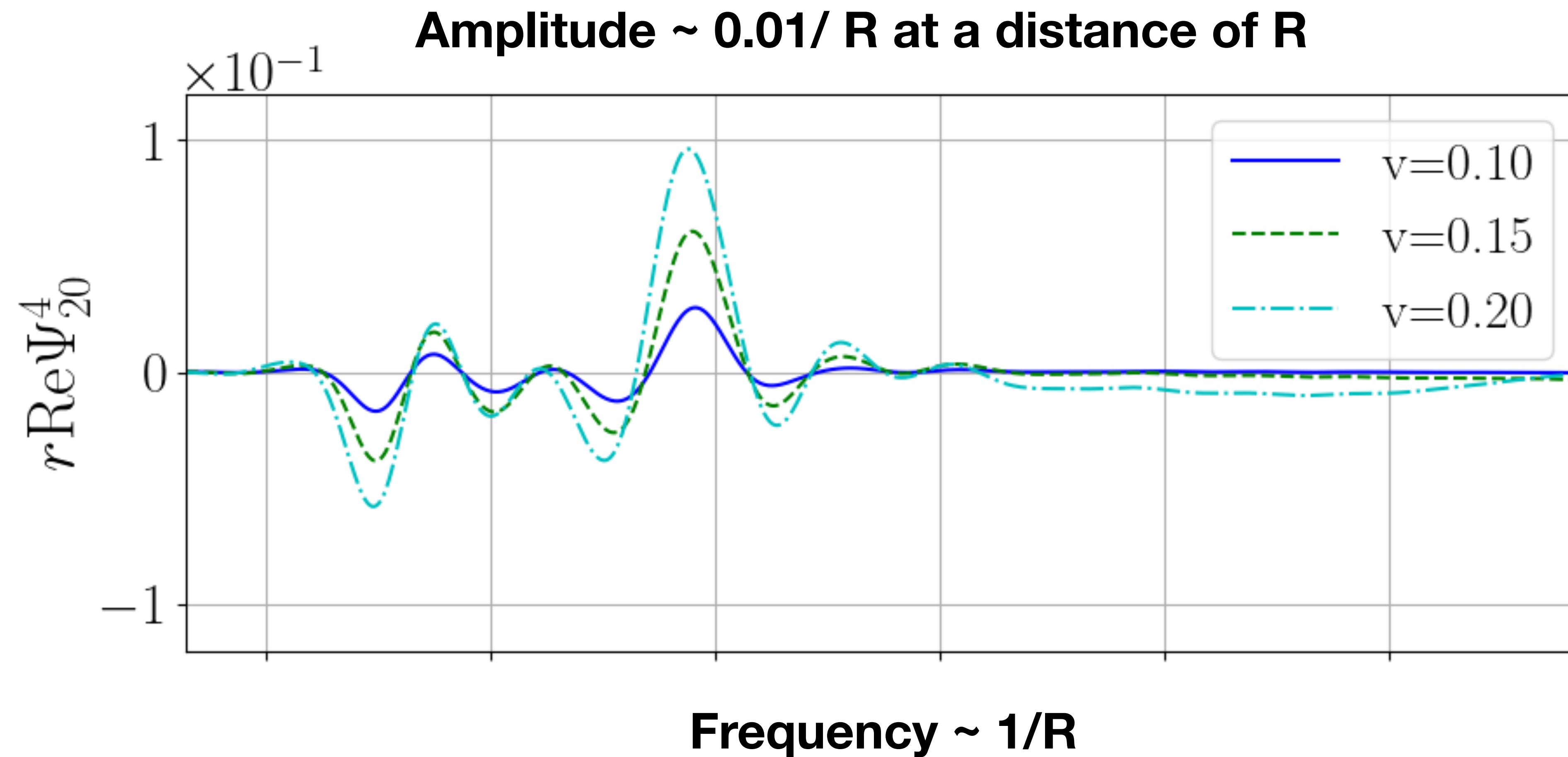
**Energy flux
of warp fluid**

Final result is a loss of negative energy, so a positive ADM mass

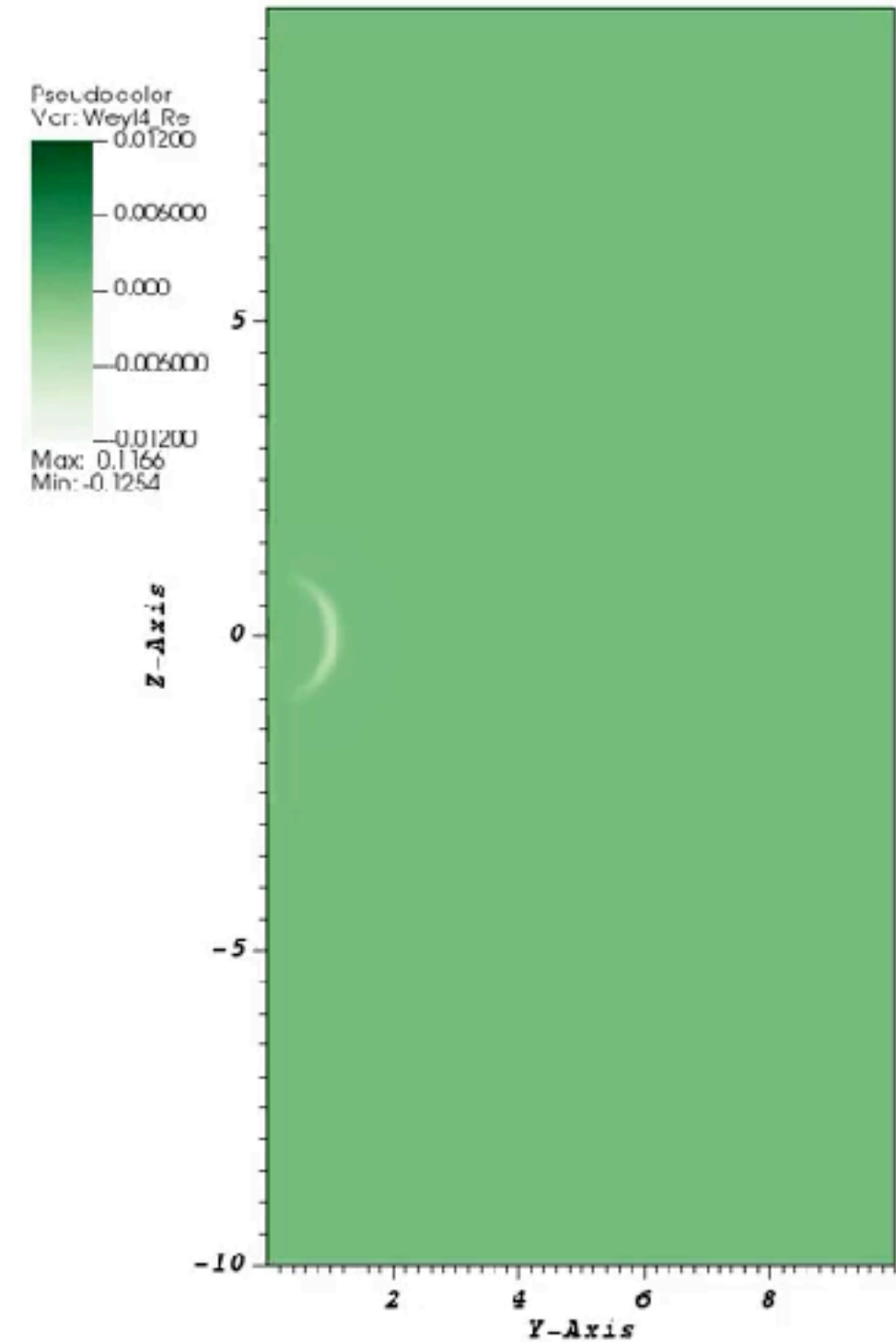
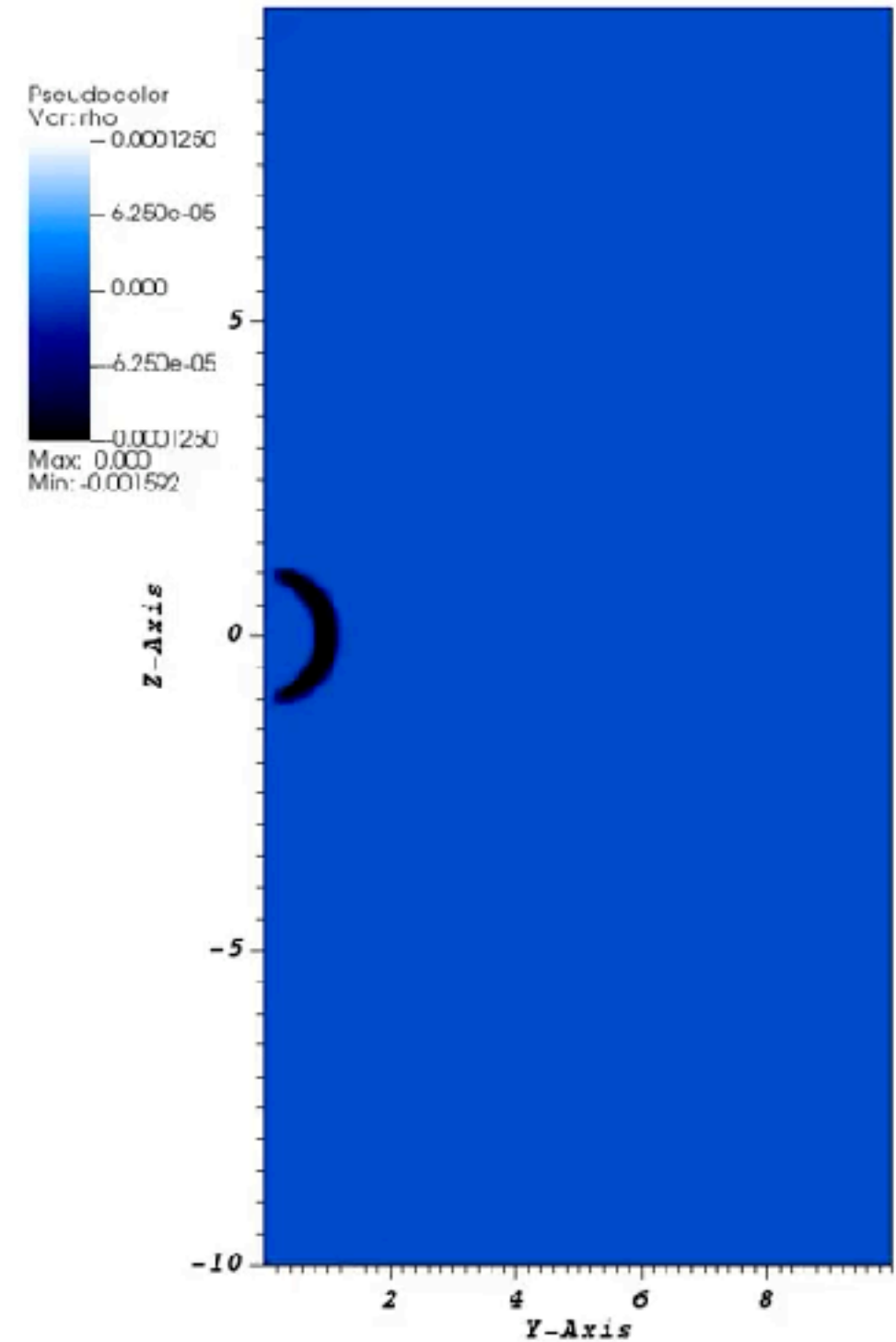
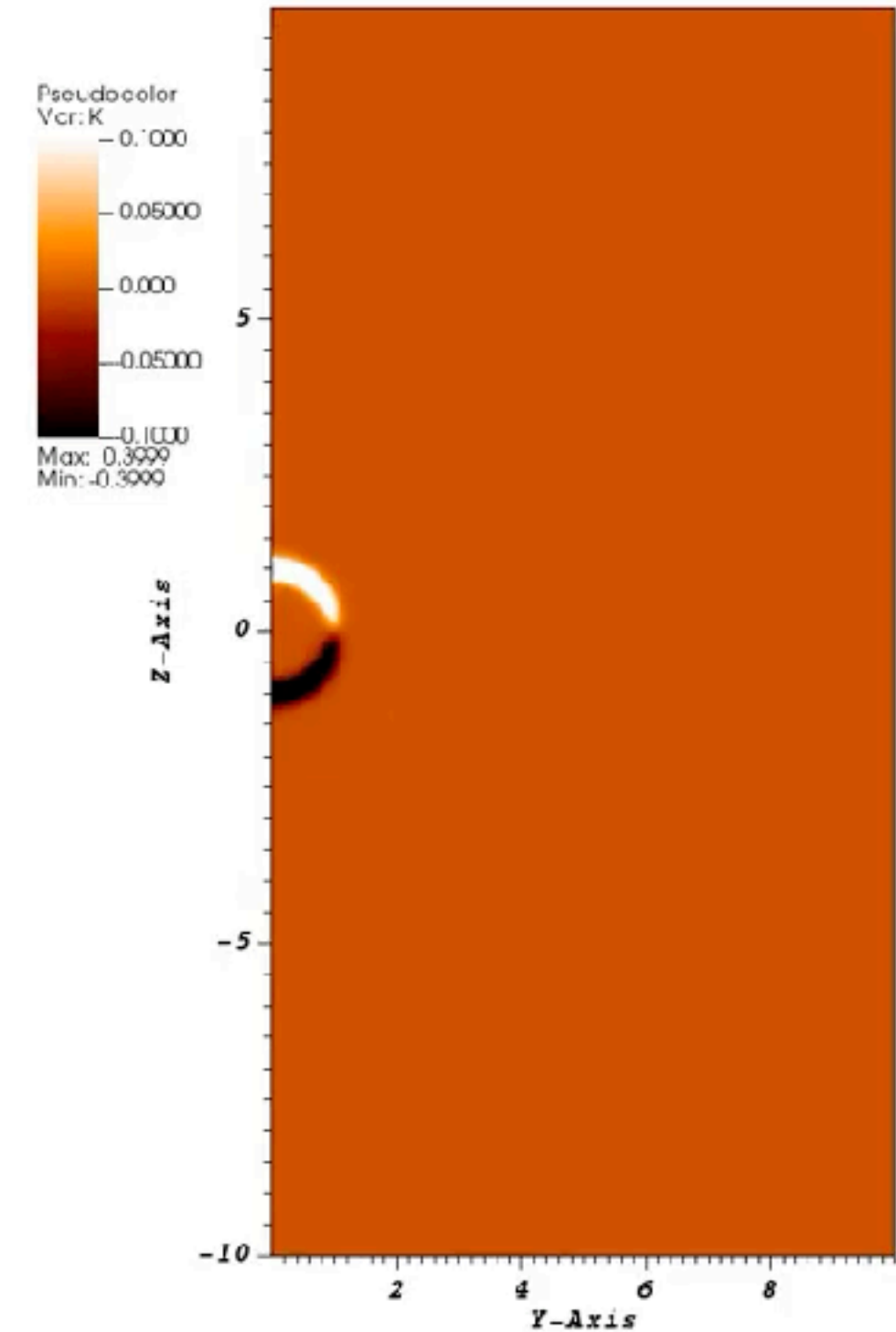


Energy fluxes

Can it be detected?

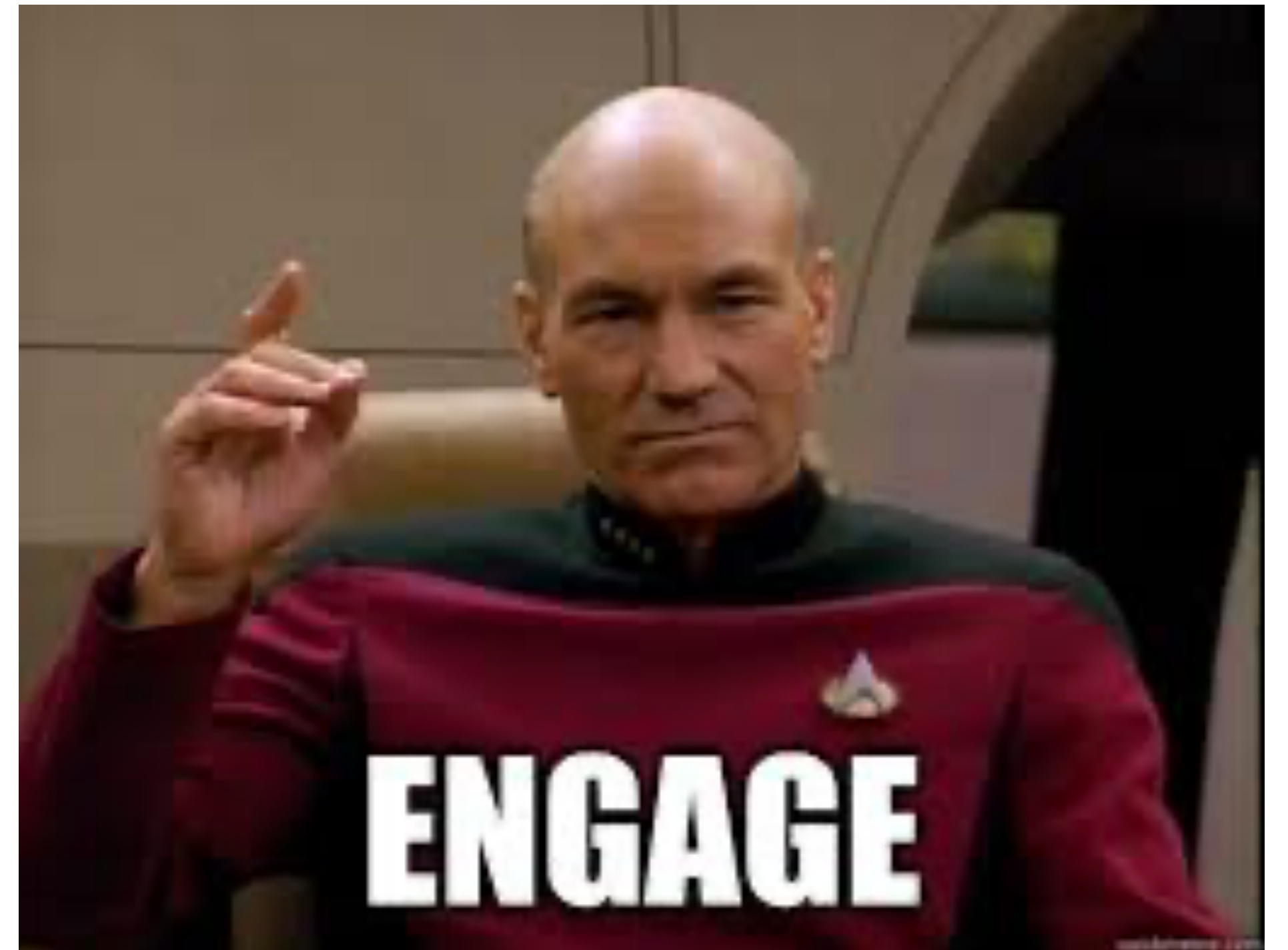


Maybe it's not the destination that matters...
it's the journey



Summary

- Don't give up on your crazy ideas (but maybe wait to get a permanent job before publishing them)
- LIGO has not ruled out post warp civilisations
- Future high frequency detectors may be able to!



Further questions

- Is this behaviour generic to negative energy spacetimes?
- What is the impact of the equation of state?
- What happens beyond $v=c$?



If you are on twitter please repost @KatyAClough

