#### Introduction to Numerical Relativity III

Erik Schnetter, Pohang, July 2007







#### Lectures Overview

- I. The Einstein Equations (Formulations and Gauge Conditions)
- II. Analysis Methods (Horizons and Gravitational Waves)
- III. Numerical methods (Cactus and Mesh Refinement)

## Numerical Methods: Cactus & Mesh Refinement

- I. Mesh Refinement
- 2. Cactus, a software framework
- 3. Kranc, a code generator
- 4. CCATIE, a free BSSN code



# Please interrupt and ask questions at any time



### Part I: Mesh Refinement



#### Mesh Refinement

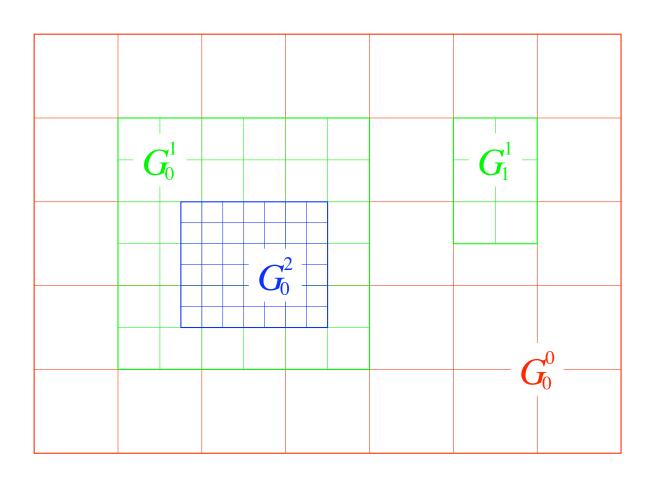
- Central idea: use different resolutions in different parts of the simulation domain
- E.g. in BBH simulations, need to cover about
   3 orders of magnitude in resolution
- Carpet: [Schnetter, Hawley, Hawke: Class.
   Quantum Grav. 21, 1465 (2004)]
- http://www.carpetcode.org/



#### Typical grid structure

Coarse, medium, and fine grids

Grids are aligned





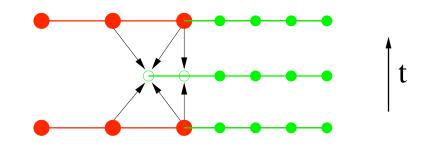
### Berger-Oliger mesh refinement

- Fine grids overlap coarse grids (coarse grids don't have holes)
- During evolution, fine grid boundary condition is interpolated from coarser grids
- Finer grids need to take smaller time steps (subcycling in time; but cf. global time stepping)
- Each grid is cuboid (rectangular), which makes it efficient

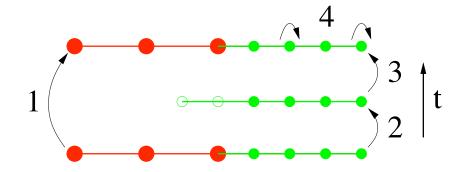


#### Time stepping

Prolongation: fine grid boundary condition



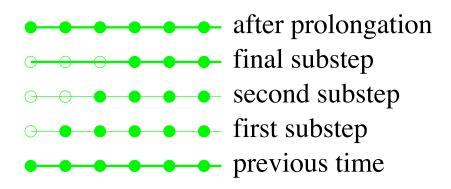
Time evolution sequence (including restriction)



Note: time interpolation requires multiple time levels



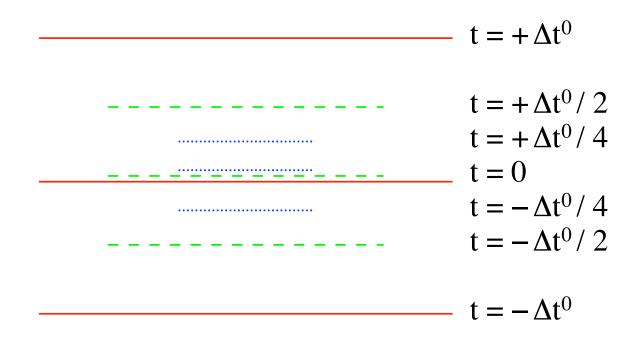
#### Buffer zones



It is unstable to prolongate during the substeps of a time integrator if there are second spatial derivatives.

Instead, use buffer zones.

#### Initial data generation



Need to set up initial data on multiple time levels. Method: evolve both forwards and backwards in time, generating an intermediate *hourglass* structure.



#### Typical usage

- Refined regions typically track black holes or neutron stars
- Typical parameters:
  - 10 levels
  - 3 ghost zones, 9 buffer zones
  - 5th order spatial, 2nd order temporal interpolation



### "Typical" simulation

- 64 processors,64 GByte memory
- Outer boundary at 500M, finest resolution 0.02M
- Total run time: several days/one week (using checkpointing and restart)



# Please interrupt and ask questions at any time



# Part II: Cactus, a software framework



#### Code Ingredients

- Evolution system, gauges (BSSN)
- Constraints, analysis quantities
- Horizon finder
- Wave extraction
- Initial data
- Test cases



#### Code Ingredients II

- Finite differencing
- Time integration
- Mesh refinement
- Parallelisation
- I/O (fast, platform-independent)
- "Something to make all this work together"



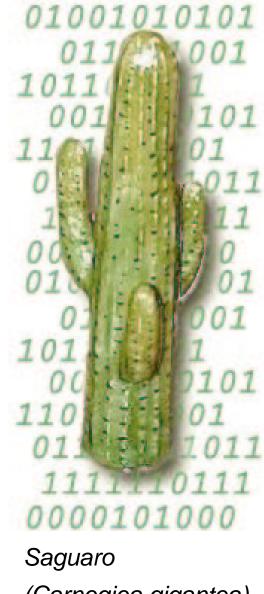
#### Components and frameworks

- Large software is typically split into components, which are large, independent pieces
- A framework puts the components together
- Cactus is a framework
- A framework doesn't do anything "useful" by itself -- it's like a bookshelf



#### Cactus

- http://www.cactuscode.org/
- In Cactus speak, the framework is called *flesh*, and the components thorns
- There are many public thorns which use Cactus, especially for numerical relativity



(Carnegiea gigantea)



#### Einstein Toolkit

- A common infrastructure for all relativity codes
- Defines common variables, common schedule events, etc.
- Comes with public thorns for basic tasks (simple initial data, simple analysis methods)

- There are least five production level relativity codes based on Cactus, all but one private, all using the Einstein Toolkit
- Three-level structure:

Physics code

Einstein Toolkit

Computational Toolkit



#### Anatomy of a Thorn

- A thorn in Cactus contains:
  - Cactus declarations (CCL language)
  - source code (C, C++, Fortran)
  - makefile fragments
  - documentation

- test cases
- example parameter files

- Thorns are the basic modular units
- Usually, each thorn is in a separate CVS repository



#### interface.ccl

- Declares thorn name and implementation name
- Declares grid functions
- Can inherit public grid functions from other implementations
- Declares routines (APIs provided/used by the thorn)

```
IMPLEMENTS: ADMConstraints
INHERITS: ADMBase

CCTK_REAL Hamiltonian TYPE=gf
{
   ham
} "Hamiltonian Constraint"

CCTK_REAL Momentum TYPE=gf
{
   momx momy momz
} "Momentum Constraint"
```



#### schedule.ccl

- Calls routines at certain times, e.g. initial or evol or analysis
- Schedule groups introduce a hierarchical structure

- Rule-based: schedule AFTER, BEFORE, WHILE
- Allocates storage for grid variables
- Synchronises variables

```
SCHEDULE ADMConstraints_Calculate AT analysis
{
   LANG: Fortran
   STORAGE: Hamiltonian Momentum
   SYNC: Hamiltonian Momentum
   TRIGGERS: Hamiltonian Momentum
} "Calculate the constraints"
```



#### param.ccl

- Declares parameters
- Five types: integer, real, boolean, keyword, string
- Allowed ranges need to be declared
- Can "inherit" public parameters from other implementations, possibly extending ranges

```
SHARES: ADMBase
EXTENTS KEYWORD initial data
  "gaussian" :: "Gaussian pulse"
}
PRIVATE:
CCTK_REAL gaussian_amplitude \
  "Amplitude"
  0.0:* :: "must be nonnegative"
} 1.0
```

#### Example Source Code

```
#include "cctk.h"
#include "cctk Arguments.h"
subroutine ADMConstraints calculate (CCTK ARGUMENTS)
  implicit none
  DECLARE CCTK_ARGUMENTS
  CCTK REAL :: dx, dy, dz
  integer :: i, j, k
  dx = CCTK DELTA SPACE(1)
  . . .
  do i = 2, cctk lsh(1)-1
    ham(i,j,k) = (gxx(i+1,j,k) - gxx(i-1,j,k)) / (2*dx)
    . . .
```



#### Parameter Files

 At run time, parameter files activate thorns and specify parameter values  Not all compiled thorns need to be active

```
ActiveThorns = "PUGH CartGrid3D ADMBase IDSimple ADMConstraints"

driver::global_nx = 101
...
grid::xmin = 0.0
grid::xmax = 30.0
...
grid::type = "octant"

ADMBase::initial_data = "Minkowski"
```



#### Driver

- A driver is a special thorn that handles memory management and parallelisation
- Two drivers exist: PUGH
   (uniform grid) and Carpet
   (AMR, multi-block)
- Two more AMR drivers in development, based on SAMRAI and Paramesh

- Interpolation, reduction, and hyperslabbing operations closely tied to driver
- I/O (efficient and parallel) and checkpointing/ recovery also somewhat driver specific



#### Metadata and Data Preservation

- Thorn Formaline collects meta-data about a simulation (and sends them to a server)
- Collects machine name, user name, parameters, current simulation time, special events, etc.

- Allows real-time overview about currently running simulations by all people on all machines
- Some simulation results are later semiautomatically staged to be permanently stored in an archive



# Please interrupt and ask questions at any time



# Part III: Kranc, a code generator



#### Coding equations

 Coding equations is very tedious; the BSSN equations contain thousands of terms

#### Example:

$$\partial_{0}\gamma_{ij} = -2\alpha K_{ij}$$

$$\partial_{t}\gamma_{ij} = -2\alpha K_{ij} + \gamma_{kj}\partial_{i}\beta^{k} + \gamma_{ik}\partial_{j}\beta^{k} + \beta^{k}\partial_{k}\gamma_{ij}$$

$$\forall_{ij} : \partial_{t}\gamma_{ij} = -2\alpha K_{ij} + \sum_{k} \gamma_{kj}\partial_{i}\beta^{k} + \sum_{k} \gamma_{ik}\partial_{j}\beta^{k} + \sum_{k} \beta^{k}\partial_{k}\gamma_{ij}$$

60 terms



#### Hand-coding

- Hand-coding takes a long time
- It is easy to make errors
- It is difficult to change the equations later
- It is also difficult to optimise the code
- Main problem: we think about the equations on a high level, but need to code on a much lower level



#### Automatic coding

- Therefore, these days most people use
   Maple or Mathematica to generate code
- E.g., Mathematica generates code fragments...
- ...and a wrapper is added by hand
- The wrapper e.g. declares variables, loads from/stores into arrays, calculates finite differences, etc.



#### Kranc

- Kranc is a Mathematica package which generates complete Cactus thorns from equations
- [Husa, Hinder, Lechner, Comput. Phys.
   Comm. I74, 983 (2006); Lechner, Alic, Husa, arXiv:cs.SC/0411063 (2004)]
- http://numrel.aei.mpg.de/Research/Kranc/



- No need to write wrappers, since whole routines including declarations are generated
- Both discretisation (finite differences) and equations are generated
- Very easy to change equations

### Disadvantages of Kranc

- Kranc adds another ingredient: not just Fortran/C code and executable, but also Mathematica script
- Need Mathematica knowledge to understand errors in Kranc script
- Less flexible than hand-coding (e.g. cannot generate HRSC formulation)



#### Demo

- I. Look at hand-written code [CactusWave]
- 2. Look a Mathematica script which uses Kranc [SW.m]
- 3. Look at Kranc-generated code [KrancSW]



# Please interrupt and ask questions at any time



# Part IV: CCATIE, a free BSSN code



#### Demo

I. Look at BBH parameter file



# Please interrupt and ask questions at any time